NUMERICAL METHODS FOR THE NAVIER–STOKES EQUATIONS. APPLICATIONS TO THE SIMULATION OF COMPRESSIBLE AND INCOMPRESSIBLE VISCOUS FLOWS

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Introduction

The main goal of this report is to discuss the numerical solution of the Navier–Stokes equations modelling steady or unsteady flows of incompressible or compressible viscous fluids. In the first part, we shall consider the incompressible case and in the second part, the compressible case. In the incompressible case, we shall particularly emphasize the time discretization by operator splitting methods, since such methods provide an efficient way to decouple the two main difficulties of the problem, namely the incompressibility and the nonlinearity. In the second part, we shall see how some of the techniques used for incompressible fluids still apply to the compressible case.

The results of numerical experiments for two-dimensional and three-dimensional flows will be presented in order to show the possibilities of the methods discussed in this report.
PART I. NUMERICAL SIMULATION OF INCOMPRESSIBLE VISCOUS FLOWS

1. Formulation of the Navier–Stokes equations for incompressible viscous fluids

Let us consider a Newtonian incompressible viscous fluid. If $\Omega$ and $\Gamma$ denote the region of the flow ($\Omega \subset \mathbb{R}^N, N = 2,3$ in practice) and its boundary, respectively, then this flow is governed by the following Navier–Stokes equations

\[ \frac{\partial u}{\partial t} - \nu \Delta u + (u \cdot \nabla) u + \nabla p = f \quad \text{in} \quad \Omega, \]  
\[ \nabla \cdot u = 0 \quad \text{in} \quad \Omega \quad \text{(incompressibility condition)}. \]  

In (1.1), (1.2):

\[ \nabla = \left( \frac{\partial}{\partial x_i} \right)_{i=1}^N, \quad \Delta = \nabla^2 = \sum_{i=1}^N \frac{\partial^2}{\partial x_i^2}, \]  
\[ u = \{ u_i \}_{i=1}^N \quad \text{is the flow velocity}, \]  
\[ p \quad \text{is the pressure}, \]  
\[ \nu \quad \text{is a viscosity parameter}, \]  
\[ f \quad \text{is a density of external forces}. \]  

In (1.1), $(u \cdot \nabla) u$ is a symbolic notation for the nonlinear vector term

\[ \left( \sum_{j=1}^N u_j \frac{\partial u_i}{\partial x_j} \right)_{i=1}^N \]

(more generally we shall denote by $(v \cdot \nabla) w$ the vector $\{ \sum_{j=1}^N v_j \frac{\partial w_i}{\partial x_j} \}_{i=1}^N$).

Boundary conditions have to be added. For example, in the case of the airfoil $A$ of fig. 1.1, we have (since the fluid is viscous) the following adherence condition

\[ u = 0 \quad \text{on} \quad \Gamma_A (= \partial A); \]  

typical conditions at infinity are

\[ u = u_\infty, \]  

where $u_\infty$ is a constant vector (with respect to the space variables, at least).

If $\Omega$ is a bounded region of $\mathbb{R}^N$, very usual boundary conditions are

\[ u = g \quad \text{on} \quad \Gamma, \]
where (from the incompressibility of the fluid) the given function \( g \) has to satisfy

\[
\int_{\Gamma} g \cdot n \, d\Gamma = 0,
\]

(1.6)

where \( n \) is the outward unit vector normal to \( \Gamma \).

Finally, for the time dependent problem (1.1), (1.2), an initial condition such as

\[
u(x, 0) = u_0(x) \quad \text{a.e. on } \Omega,
\]

(1.7)

with \( u_0 \) given, is usually prescribed.

In practice, for the problem corresponding to fig. 1.1, we should replace \( \Omega \) by a large bounded domain \( \Omega_c \) (the computational domain), and on the external boundary \( \Gamma_\infty \) of \( \Omega_c \) we should prescribe \( u = u_\infty \), or some more sophisticated boundary conditions.

**Remark 1.1:** For two-dimensional problems on unbounded domains \( \Omega \), exponential stretching methods can be used, allowing very large computational domains (see ref. [1], chap. 7 and ref. [2] for applications of exponential stretching methods to inviscid flow calculations).

**Remark 1.2:** When using \( \Omega_c \) (as above) instead of \( \Omega \), prescribing \( u = u_\infty \) on the whole \( \Gamma_\infty \) may not be satisfactory if \( \Omega_c \) is not sufficiently large. Actually, we should improve the computed solutions using as boundary conditions

\[
u = u_\infty \quad \text{on } \Gamma_\infty^-,
\]

(1.8)

and either

\[
\nu \frac{\partial u}{\partial n} - n_\infty p = 0 \quad \text{on } \Gamma_\infty^+,
\]

(1.9a)
or
\[
\frac{\partial}{\partial t} \omega + c \frac{\partial \omega}{\partial n} = 0 \quad \text{on} \quad \Gamma^+_\infty,
\]

where
1. \( \Gamma^+_\infty = \{ x \mid x \in \Gamma_\infty, \ u_\infty \cdot n_\infty(x) > 0 \} \),
2. \( n_\infty \) is the outward unit vector normal to \( \Gamma_\infty \) (see fig. 1.1),
3. \( c \) is a constant; a natural choice seems to be \( c = |u_\infty| \),
4. \( \omega = \nabla \times u \) is the vorticity of the flow.

The main reason for using either (1.8), (1.9a) or (1.9b), instead of \( u = u_\infty \) on \( \Gamma_\infty \), is that the former boundary conditions are less reflecting (i.e. more absorbing) than the latter.

The numerical implementation of more sophisticated absorbing boundary conditions, taking advantage of the time discretization and of the finite element approximation, will be discussed in section 7.

The Navier–Stokes equations for incompressible viscous fluids have motivated a large number of papers, books, reports and symposia; we shall limit our references to:
1. refs. [3–8] for the theoretical aspects,
2. refs. [1,6,8–12] for the numerical aspects.

In both cases, the interested reader may consult the references contained in the above books.

The difficulties with the Navier–Stokes equations (even for flows at low Reynold’s numbers, in bounded regions \( \Omega \)) are
1. The nonlinear term \( (u \cdot \nabla)u \) in (1.1);
2. The incompressibility condition (1.2);
3. The fact that the solutions of the Navier–Stokes equations are vector-valued functions of \( x, t \), whose components are coupled by the nonlinear term \( (u \cdot \nabla)u \) and by the incompressibility condition \( \nabla \cdot u = 0 \).

Using convenient operator splitting methods for the time discretization of the Navier–Stokes equations, we shall be able to decouple those difficulties associated to the nonlinearity and to the incompressibility, respectively.

For simplicity, we suppose from now on that \( \Omega \) is bounded and that we have (1.5) as boundary conditions (with \( g \) satisfying (1.6) and possibly depending upon \( t \)).

2. Time discretization by operator splitting methods

2.1. Generalities. Description of the basic schemes

Let us consider a real Hilbert space \( H \); we consider then in \( H \) the following initial value problem

\[
\begin{aligned}
\frac{du}{dt} + A(u) &= f, \\
u(0) &= u_0,
\end{aligned}
\]
where \( A \) is an operator from \( H \) to \( H \) and where \( f \) is a source term and \( u_0 \) the initial value, respectively. Let \( A_1 \) and \( A_2 \) be two operators such that

\[
A = A_1 + A_2. \tag{2.2}
\]

With \( \Delta t > 0 \) a time discretization step, let us define several schemes for solving (2.1), taking advantage of the decomposition (2.2):

**A first scheme (Peaceman–Rachford type):** The scheme is defined as follows

\[
u^0 = u_0; \tag{2.3}
\]

then for \( n \geq 0 \), with \( u^n \) known, compute \( u^{n+1/2} \) and then \( u^{n+1} \) by

\[
\begin{align*}
\frac{u^{n+1/2} - u^n}{\Delta t/2} + A_1(u^{n+1/2}) + A_2(u^n) &= f^{n+1/2}, \\
\frac{u^{n+1} - u^{n+1/2}}{\Delta t/2} + A_1(u^{n+1/2}) + A_2(u^{n+1}) &= f^{n+1}.
\end{align*} \tag{2.4, 2.5}
\]

In (2.4), (2.5), \( u^{n+\alpha} \) denotes an approximation of \( u((n + \alpha)\Delta t) \), and \( f^{n+\alpha} = f((n + \alpha)\Delta t) \).

**A second scheme (\( \theta \)-scheme):** Let \( \theta \) belong to the open interval \((0, 1/2)\). The idea behind the scheme is to split the time interval \([n\Delta t, (n + 1)\Delta t]\) into three subintervals, as shown in fig. 2.1, and integrate (approximately) with respect to time (using an implicit scheme for \( A_1 \) and an explicit scheme for \( A_2 \)) on \([n\Delta t, (n + \theta)\Delta t]\), then switch the role of \( A_1 \) and \( A_2 \) on \([(n + \theta)\Delta t, (n + 1 - \theta)\Delta t]\), and finally, on \([(n + 1 - \theta)\Delta t, (n + 1)\Delta t]\) do like on \([n\Delta t, (n + \theta)\Delta t]\).

Using these principles we obtain the following scheme, some forms of which have been advocated in refs. [13–16] (for \( \theta = 1/4 \)):

\[
u^0 = u_0; \tag{2.6}
\]

then, for \( n \geq 0 \), we obtain \( u^{n+\theta}, u^{n+1-\theta}, u^{n+1} \), from \( u^n \), as follows

\[
\begin{align*}
\frac{u^{n+\theta} - u^n}{\theta \Delta t} + A_1(u^{n+\theta}) + A_2(u^n) &= f^{n+\theta}, \\
\frac{u^{n+1-\theta} - u^{n+\theta}}{(1 - 2\theta) \Delta t} + A_1(u^{n+\theta}) + A_2(u^{n+1-\theta}) &= f^{n+1-\theta}, \\
\frac{u^{n+1} - u^{n+1-\theta}}{\theta \Delta t} + A_1(u^{n+1}) + A_2(u^{n+1-\theta}) &= f^{n+1}. \quad \square
\end{align*} \tag{2.7, 2.8, 2.9}
\]

![Fig. 2.1.](image-url)
The convergence of (2.3)–(2.5) has been proved in ref. [17] (see also ref. [18]) under quite general monotonicity assumptions on \( A_1 \) and \( A_2 \); we don’t know similar general results for (2.6)–(2.9) (but, very likely, the methods in ref. [17] still apply).

2.2. Convergence and stability properties of the basic schemes

Following the approach used in ref. [19], we consider for simplicity the case where \( H = \mathbb{R}^N \), \( f = 0 \), \( u_0 \in \mathbb{R}^N \), \( A \) is an \( N \times N \) symmetric and positive definite matrix and where

\[
A_1 = \alpha A, \quad A_2 = \beta A, \quad \text{with} \quad \alpha + \beta = 1, \quad 0 < \alpha, \beta < 1.
\]

In that case, the solution of (2.1) is clearly given by

\[
u(t) = e^{-\alpha t} u_0.
\]

Analysis of scheme (2.3)–(2.5) (see also ref. [19]): We have, from (2.4), (2.5), (2.10)

\[
u^{n+1} = (I + \beta (\Delta t/2) A)^{-1}(I - \alpha (\Delta t/2) A)(I + \alpha (\Delta t/2) A)^{-1}(I - \beta (\Delta t/2) A) u^n.
\]

Using a vector basis consisting of eigenvectors of \( A \), we have, from (2.12), with obvious notation,

\[
u^{n+1} = \frac{(1 - \alpha (\Delta t/2) \lambda_i)(1 - \beta (\Delta t/2) \lambda_i)}{(1 + \alpha (\Delta t/2) \lambda_i)(1 + \beta (\Delta t/2) \lambda_i)} u^n.
\]

where \( \lambda_i (> 0, \forall i = 1, \ldots, N) \) is the \( i \)th eigenvalue of \( A \); we suppose that \( \lambda_1 \leq \lambda_2 \leq \cdots \leq \lambda_N \).

Consider now the rational function \( R_1 \) defined by

\[
R_1(x) = \frac{(1 - (\alpha/2)x)(1 - (\beta/2)x)}{(1 + (\alpha/2)x)(1 + (\beta/2)x)};
\]

we observe that \( |R_1(x)| < 1, \forall x > 0 \), implying, in that simple case, the unconditional stability of scheme (2.3)–(2.5). Since

\[
\lim_{x \to +\infty} R_1(x) = 1,
\]

we observe that for stiff problems, i.e. problems such that \( \lambda_N/\lambda_1 \gg 1 \), scheme (2.3)–(2.5) is not very good to damp simultaneously, the components of \( u^n \) associated to the large and to the small eigenvalues of \( A \). From this observation, we can expect that scheme (2.3)–(2.5) is not well-suited to “capture” the steady state solutions of stiff problems (like those obtained from the discretization of partial differential equations); this has been confirmed by numerical experiments. Let us discuss now the accuracy of scheme (2.3)–(2.5). Since

\[
e^{-x} = 1 - x + \frac{1}{2}x^2 + x^2 \epsilon(x),
\]
and, from (2.14),

$$R_1(x) = 1 - x + \frac{1}{2}x^2 + x^2\eta(x),$$  \hspace{1cm} (2.17)

with \(\lim_{x \to 0} \eta(x) = 0\), we have that scheme (2.3)–(2.5) is second order accurate in the simple case considered here. We observe, from (2.12), that if one takes \(\alpha = \beta = \frac{1}{2}\), then the two linear systems which have to be solved, at each full step, are, in fact, associated to the same matrix \(I + (\Delta t/4)A\).

**Analysis of scheme (2.6)–(2.9):**

We have (with \(\theta' = 1 - 2\theta\))

$$\begin{align*}
\begin{cases}
  u^{n+1} = & (I + \alpha_\theta \Delta tA)^{-1}(I - \beta_\theta \Delta tA)(I + \beta_\theta' \Delta tA)^{-1}(I - \alpha_\theta' \Delta tA) \\
  & \times (I + \alpha_\theta \Delta tA)^{-1}(I - \beta_\theta \Delta tA)u^n,
\end{cases}
\end{align*}$$

(2.18)

which implies

$$u_i^{n+1} = \frac{(1 - \beta_\theta \Delta t\lambda_i)^2(1 - \alpha_\theta' \Delta t\lambda_i)}{(1 + \alpha_\theta \Delta t\lambda_i)^2(1 + \beta_\theta' \Delta t\lambda_i)} u_i^n.$$  \hspace{1cm} (2.19)

Consider now the rational function \(R_2\) defined by

$$R_2(x) = \frac{(1 - \beta_\theta x)^2(1 - \alpha_\theta' x)}{(1 + \alpha_\theta x)^2(1 + \beta_\theta' x)}.$$  \hspace{1cm} (2.20)

Since

$$\lim_{x \to +\infty} |R_2(x)| = \beta/\alpha,$$  \hspace{1cm} (2.21)

we should prescribe

$$\alpha > \beta,$$  \hspace{1cm} (2.22)

to have, from (2.18), (2.19), the stability of scheme (2.6)–(2.9) for the large eigenvalues of \(A\). We discuss now the accuracy of scheme (2.6)–(2.9); we can show that

$$R_2(x) = 1 - x + \frac{1}{2}x^2\left(1 + (\beta^2 - \alpha^2)(2\theta^2 - 4\theta + 1)\right) + x^2\eta(x),$$  \hspace{1cm} (2.23)

with \(\lim_{x \to 0} \eta(x) = 0\). It follows from (2.23) that scheme (2.6)–(2.9) is second order accurate if, either

$$\alpha = \beta (= \frac{1}{2} \text{ from (2.10))},$$  \hspace{1cm} (2.24)
or

$$\theta = 1 - \sqrt{2} / 2 = 0.29289 \ldots ;$$  \hspace{1cm} (2.25)

scheme (2.6)–(2.9) is only first order accurate if neither (2.24) nor (2.25) hold. If one takes
\( \alpha = \beta = \frac{1}{2} \), it follows from (2.19) that scheme (2.6)–(2.9) is unconditionally stable, \( \forall \theta \in ]0, \frac{1}{2}[ \); however, since (from (2.21)) we have

$$\lim_{x \to +\infty} | R_2(x) | = 1,$$  \hspace{1cm} (2.26)

the remark stated for scheme (2.3)–(2.5) concerning the integration of stiff systems still holds. In
general, we shall choose \( \alpha \) and \( \beta \) in order to have the same matrix for all the partial steps of the
integration procedure; i.e., \( \alpha, \beta, \theta \) have to satisfy

$$\alpha \theta = \beta (1 - 2 \theta),$$  \hspace{1cm} (2.27)

which implies

$$\alpha = (1 - 2 \theta) / (1 - \theta), \quad \beta = \theta / (1 - \theta).$$  \hspace{1cm} (2.28)

Combining (2.22), (2.28) we obtain

$$0 < \theta \leq \frac{1}{4}.$$  \hspace{1cm} (2.29)

For \( \theta = \frac{1}{4} \), (2.28) implies \( \alpha = \beta = \frac{1}{2} \); the resulting scheme is just a variant of scheme (2.3)–(2.5).

If \( 0 < \theta < \frac{1}{4} \), and if \( \alpha \) and \( \beta \) are given by (2.28), we have then

$$\lim_{x \to +\infty} | R_2(x) | = \beta / \alpha = \theta / (1 - 2 \theta) < 1.$$  \hspace{1cm} (2.30)

Actually, we can prove that \( \theta \in [\theta^*, \frac{1}{4}] \) (with \( \theta^* = 0.087385580 \ldots \)) and \( \alpha, \beta \) given by (2.28),

imply the unconditional stability of scheme (2.6)–(2.9). Moreover, if \( \theta \in [\theta^*, \frac{1}{4}] \), property (2.30)

makes scheme (2.6)–(2.9) have good asymptotic properties as \( n \to +\infty \) and for example, is well

suited to compute steady state solutions. If \( \theta = 1 - \sqrt{2} / 2 \) (resp. \( \theta = \frac{1}{4} \)), we have \( \alpha = 2 - \sqrt{2}, \beta = \sqrt{2} - 1, \beta / \alpha = 1 / \sqrt{2} \) (resp. \( \alpha = \frac{1}{8}, \beta = \frac{1}{8}, \beta / \alpha = \frac{1}{8} \)).

2.3. Application to the solution of the time dependent Navier–Stokes equations

2.3.1. A first operator splitting method

This method is directly derived from the Peaceman–Rachford scheme (2.3)–(2.5) and is

described as follows:

$$u^o = u_0,$$  \hspace{1cm} (2.31)
then for \( n \geq 0 \), compute, from \( u^n \), \( \{ u^{n+1/2}, p^{n+1/2} \} \) and \( u^{n+1} \) by solving

\[
\begin{align*}
\left\{
\begin{array}{l}
\frac{u^{n+1/2} - u^n}{\Delta t} - \frac{\nu}{2} \Delta u^{n+1/2} + \nabla p^{n+1/2} &= f^{n+1/2} + \frac{\nu}{2} \Delta u^n - (u^n \cdot \nabla) u^n \quad \text{in} \ \Omega, \\
\nabla \cdot u^{n+1/2} &= 0 \quad \text{in} \ \Omega, \\
u p_n^{n+1/2} &= g_n^{n+1/2} \quad \text{on} \ \Gamma,
\end{array}
\right.
\end{align*}
\]

(2.32)

\[
\begin{align*}
\left\{
\begin{array}{l}
\frac{u^{n+1} - u^n}{\Delta t} - \frac{\nu}{2} \Delta u^{n+1} + (u^{n+1} \cdot \nabla) u^{n+1} - f^{n+1} + \frac{\nu}{2} \Delta u^{n+1/2} - \nabla p^{n+1/2} \\
in \ \Omega, \\
u p_n^{n+1} &= g_n^{n+1} \quad \text{on} \ \Gamma,
\end{array}
\right.
\end{align*}
\]

(2.33)

respectively.

2.3.2. A second operator splitting method

This method is derived from scheme (2.6)-(2.9) and is obtained as follows:

\[
u_0 = u_0,
\]

(2.34)

then for \( n \geq 0 \), starting from \( u^n \) we solve

\[
\begin{align*}
\left\{
\begin{array}{l}
\frac{u^{n+\theta} - u^n}{\theta} - \frac{\nu}{\Delta t} \Delta u^{n+\theta} + \nabla p^{n+\theta} &= f^{n+\theta} + \beta \nu \Delta u^n - (u^n \cdot \nabla) u^n \quad \text{in} \ \Omega, \\
\nabla \cdot u^{n+\theta} &= 0 \quad \text{in} \ \Omega, \\
u p_n^{n+\theta} &= g_n^{n+\theta} \quad \text{on} \ \Gamma,
\end{array}
\right.
\end{align*}
\]

(2.35)

\[
\begin{align*}
\left\{
\begin{array}{l}
\frac{u^{n+1-\theta} - u^{n+\theta}}{(1 - 2\theta) \Delta t} - \beta \nu \Delta u^{n+1-\theta} + (u^{n+1-\theta} \cdot \nabla) u^{n+1-\theta} \\
= f^{n+1-\theta} + \alpha \nu \Delta u^{n+\theta} - \nabla p^{n+\theta} \quad \text{in} \ \Omega, \\
u p_n^{n+1} &= g_n^{n+1} \quad \text{on} \ \Gamma,
\end{array}
\right.
\end{align*}
\]

(2.36)

\[
\begin{align*}
\left\{
\begin{array}{l}
\frac{u^{n+1} - u^{n+1-\theta}}{\theta \Delta t} - \frac{\nu}{\Delta t} \Delta u^{n+1} + \nabla p^{n+1} \\
= f^{n+1} + \beta \nu \Delta u^{n+1-\theta} - (u^{n+1-\theta} \cdot \nabla) u^{n+1-\theta} \quad \text{in} \ \Omega, \\
\nabla \cdot u^{n+1} &= 0 \quad \text{in} \ \Omega, \\
u p_n^{n+1} &= g_n^{n+1} \quad \text{on} \ \Gamma.
\end{array}
\right.
\end{align*}
\]

(2.37)

2.3.3. Some comments and remarks concerning schemes (2.31)-(2.33) and (2.34)-(2.37)

Using the two above operator splitting methods, we have been able to decouple nonlinearity and incompressibility in the Navier–Stokes eqs. (1.1), (1.2). We shall describe, in the following
sections, the specific treatment of the subproblems encountered at each step of (2.31)-(2.33) and (2.34)-(2.37). We shall only consider the case where the subproblems are still continuous in space (since the formalism of the continuous problems is much simpler); for the fully discrete case, see refs. [1,20] where finite element approximations of (1.1), (1.2) are discussed.

We observe that \( u^{n+1/2} \) and \( u^{n+\theta}, u^{n+1} \) are obtained from the solutions of linear problems very close to the steady Stokes problem. Despite its greater complexity, scheme (2.30)-(2.37) is almost as economical to use as scheme (2.31) (2.33); this is mainly due to the fact that the “quasi” steady Stokes problems (2.32) and (2.35), (2.37) (actually convenient finite element approximations of them) can be solved by quite efficient solvers, so that most of the computer time used to solve a full step is, in fact, used to solve the nonlinear subproblem.

If one uses scheme (2.34)-(2.37), the best choice for \( \alpha \) and \( \beta \) is given by (2.28). With such a choice, many computer subprograms can be used for both the linear and nonlinear subproblems, resulting, therefore, in quite substantial core memory savings.

We can find, in ref. [9, chap. 4], a variant of scheme (2.31)-(2.33), given by

\[ u^0 = u_0; \]

then for \( n \geq 0 \), starting from \( u^n \), compute \( \{ u^{n+1/2}, p^{n+1/2} \} \) and \( u^{n+1} \) by

\[
\begin{cases}
\frac{u^{n+1/2} - u^n}{\Delta t/2} - \nu \Delta u^{n+1/2} + \nabla p^{n+1/2} = f^{n+1/2} - (u^n \cdot \nabla) u^n \quad &\text{in } \Omega, \\
\nabla \cdot u^{n+1/2} = 0 \quad &\text{in } \Omega, \\
u^{n+1/2} = g^{n+1/2} \quad &\text{on } \Gamma, 
\end{cases}
\]

and

\[
\begin{cases}
\frac{u^{n+1} - u^{n+1/2}}{\Delta t/2} - \nu \Delta u^n + (u^{n+1} \cdot \nabla) u^{n+1} = f^{n+1} - \nabla p^{n+1/2} \quad &\text{in } \Omega, \\
u^{n+1} = g^{n+1} \quad &\text{on } \Gamma.
\end{cases}
\]

3. Least squares–conjugate gradient solution of the nonlinear subproblems

3.1. Classical and variational formulations. Synopsis

At each full step of the operator splitting methods (2.31)-(2.33) and (2.34)-(2.37), we have to solve a nonlinear elliptic system of the following type

\[
\begin{cases}
\alpha u - \nu \Delta u + (u \cdot \nabla) u = f \quad &\text{in } \Omega, \\
u = g \quad &\text{on } \Gamma,
\end{cases}
\]

where \( \alpha \) and \( \nu \) are two positive parameters and where \( f \) and \( g \) are two given functions defined
on $\Omega$ and $\Gamma$, respectively. We do not discuss here the existence and uniqueness of solutions for problem (3.1).

We introduce now the following functional spaces of Sobolev’s type (see e.g. refs. [21–24] for information on Sobolev spaces):

\[
H^1(\Omega) = \{ \phi | \phi \in L^2(\Omega), \; \partial \phi / \partial x_i \in L^2(\Omega), \; \forall i = 1, \ldots, N \},
\]

(3.2)

\[
H_0^1(\Omega) = \{ \phi | \phi \in H^1(\Omega), \; \phi = 0 \; \text{on} \; \Gamma \},
\]

(3.3)

\[
V_0 = (H_0^1(\Omega))^N,
\]

(3.4)

\[
V_g = \{ v | v \in (H^1(\Omega))^N, \; v = g \; \text{on} \; \Gamma \};
\]

(3.5)

if $g$ is sufficiently smooth, then $V_g$ is nonempty.

We shall use the following notation

\[
dx = dx_1 \, dx_2 \cdots dx_N,
\]

and if

\[
u = \{ u_i \}_{i=1}^N, \quad v = \{ v_i \}_{i=1}^N,
\]

then

\[
u \cdot v = \sum_{i=1}^N u_i v_i,
\]

\[
\nabla \nu \cdot \nabla v = \sum_{i=1}^N \nabla u_i \cdot \nabla v_i = \sum_{i=1}^N \sum_{j=1}^N \frac{\partial u_i}{\partial x_j} \frac{\partial v_i}{\partial x_j}.
\]

Using Green’s formula we can prove that for sufficiently smooth functions $\nu$ and $v$, belonging to $(H^1(\Omega))^N$ and $V_0$, respectively, we have

\[
- \int_{\Omega} \Delta \nu \cdot v \, dx = \int_{\Omega} \nabla \nu \cdot \nabla v \, dx.
\]

(3.6)

It can also be shown that $\nu$ is a solution of the nonlinear variational problem

\[
\begin{cases}
\nu \in V_g, \\
\alpha \int_{\Omega} \nu \cdot v \, dx + \int_{\Omega} \nabla \nu \cdot \nabla v \, dx + \int_{\Omega} (\nu \cdot \nabla) \nu \cdot v \, dx = \int_{\Omega} f \cdot v \, dx,
\end{cases}
\]

(3.7)

and conversely. We observe that (3.1), (3.7) is not equivalent to a problem of the Calculus of Variations since there is no functional of $v$ with $(v \cdot \nabla)v$ as differential. However, using a
convenient least-squares formulation, we shall be able to solve (3.1), (3.7) by iterative methods originating from Nonlinear Programming, such as Conjugate Gradient for example. Indeed, since problem (3.1), (3.7) is nonlinear, a first natural choice for solving it is Newton's Method; such a method is described in section 3.2, just below.

3.2. On the solution of (3.1), (3.7) by Newton's method

We recall first some general principles concerning Newton's method:
Let X and Y be two real Banach spaces and $F$ a $C^1$ (possibly nonlinear) operator from X into Y. The Newton's method for solving the following equation

\[ F(u) = 0 \]  

is defined by

\[ u^0 \in X, \text{ given}; \]  

then for $m \geq 0$, we define $u^{m+1}$ from $u^m$ by

\[ u^{m+1} = u^m - \left( F'(u^m) \right)^{-1} F(u^m), \]  

where $F'(u^m) \in \mathbb{L}(X, Y)$ is the derivative of $F$ at $u^m$.

Applied to the nonlinear problem (3.1), (3.7), the Newton's algorithm (3.9), (3.10) yields

\[ u^0 \in V_g, \text{ given}; \]  

then for $m \geq 0$, compute $u^{m+1}$ from $u^m$, by solving the following linear elliptic problem

\[
\begin{aligned}
\alpha u^{m+1} - \nu \Delta u^{m+1} + (u^{m+1} \cdot \nabla) u^m + (u^m \cdot \nabla) u^{m+1} &= (u^m \cdot \nabla) u^m + f \\
&\text{in } \Omega,
\end{aligned}
\]  

\[ u^{m+1} = g \text{ on } \Gamma, \]  

whose variational formulation is given by

\[
\begin{aligned}
&u^{m+1} \in V_g, \quad \forall v \in V_0, \\
&\alpha \int_{\Omega} u^{m+1} \cdot v \, dx + \nu \int_{\Omega} \nabla u^{m+1} \cdot \nabla v \, dx + \int_{\Omega} (u^{m+1} \cdot \nabla) u^m \cdot v \, dx + \int_{\Omega} (u^m \cdot \nabla) u^{m+1} \cdot v \, dx = \int_{\Omega} (u^m \cdot \nabla) u^m \cdot v \, dx + \int_{\Omega} f \cdot v \, dx.
\end{aligned}
\]

Problem (3.12), (3.13) is associated to a nonself-adjoint elliptic operator, varying with $m$. After an appropriate approximation, we should obtain a linear system associated to a nonsymmetric matrix, varying with $m$. Solving such a linear system may be quite complicated and costly and it
is for that reason that we have developed alternative solution methods, founded on least squares formulations and conjugate gradient algorithms. Such methods are described in sections 3.3, 3.4, below.

### 3.3. Least squares formulation of the nonlinear problem (3.1), (3.7)

Let $v \in V_g$; from $v$, we define $y(-y(v)) \in V_0$ as the solution of

\[
\begin{cases}
\alpha y - \nu \Delta y = \alpha v - \nu \Delta v + (v \cdot \nabla) v - f & \text{in } \Omega, \\
y = 0 & \text{on } \Gamma.
\end{cases}
\] (3.14)

We observe that $y$ is obtained from $v$ via the solution of $N$ uncoupled linear Poisson problems (one for each component of $y$); using (3.6) it can be shown that problem (3.14) is actually equivalent to the linear variational problem

\[
\begin{align*}
\text{Find } & y \in V_0 \text{ such that, } \forall z \in V_0 \text{ we have } \\
& \int_{\Omega} (\alpha y \cdot z + \nu \int_{\Omega} \nabla y \cdot \nabla z \, dx \\
& = \alpha \int_{\Omega} v \cdot z \, dx + \nu \int_{\Omega} \nabla v \cdot \nabla z \, dx + \int_{\Omega} (v \cdot \nabla) v \cdot z \, dx - \int_{\Omega} f \cdot z \, dx,
\end{align*}
\] (3.15)

which has a unique solution. Suppose now that $v$ is a solution of the nonlinear problem (3.1), (3.7); the corresponding $y$ (obtained from the solution of (3.14), (3.15)) is clearly $y = 0$. From this observation, it is quite natural to introduce the following (nonlinear) least squares formulation of (3.1), (3.7)

\[
\begin{cases}
\text{Find } u \in V_g \text{ such that } \\
J(u) \leq J(v), \quad \forall v \in V_g,
\end{cases}
\] (3.16)

where $J: (H^1(\Omega))^N \to \mathbb{R}$ is the function of $v$ defined by

\[
J(v) = \frac{1}{2} \int_{\Omega} \{ \alpha | y |^2 + \nu | \nabla y |^2 \} \, dx,
\] (3.17)

and where $y$ is defined from $v$ by (3.14), (3.15). We observe that if $u$ is a solution of (3.1), (3.7), then it is also a solution of (3.16), such that $J(u) = 0$. Conversely, if $u$ is a solution of (3.16) such that $J(u) = 0$, then it is also a solution of (3.1), (3.7).

We conclude this paragraph by the computation of $J'(v)$ where $J'$ is the derivative of $J$. Let $\delta v$ be a "perturbation" of $v$ compatible with the boundary condition $v = g$ on $\Gamma$. We have then $\delta v = 0$ on $\Gamma$, implying that $\delta v \in V_0$. We have, from (3.17),

\[
\delta J = \langle J'(v), \delta v \rangle = \int_{\Omega} \{ \alpha y \cdot \delta y + \nu \nabla y \cdot \nabla \delta y \} \, dx.
\] (3.18)
We also have, from (3.15),
\[
\begin{aligned}
\frac{\delta y}{\delta y} \in V_0; & \quad \forall z \in V_0, \quad \text{we have} \\
\alpha \int_\Omega \delta y \cdot \nabla z \, dx + \nu \int_\Omega \nabla \delta y \cdot \nabla z \, dx \\
& = \alpha \int_\Omega \delta y \cdot z \, dx + \nu \int_\Omega \nabla \delta y \cdot \nabla z \, dx + \int_\Omega \int (\delta v \cdot \nabla) v \cdot z \, dx + \int_\Omega (\delta v \cdot \nu) \delta v \cdot z \, dx.
\end{aligned}
\] (3.19)

Taking \( z = y \) in (3.19), it follows from (3.18) that
\[
\langle J'(v), \delta v \rangle = \alpha \int_\Omega y \cdot \delta v \, dx + \nu \int_\Omega \nabla y \cdot \nabla \delta v \, dx + \int_\Omega y \cdot (\delta v \cdot \nabla) v \, dx + \int_\Omega y \cdot (v \cdot \nabla) \delta v \, dx,
\]
which implies in turn that \( J'(v) \) can be identified with the linear functional from \( V_0 \) to \( \mathbb{R} \), defined by
\[
\langle J'(v), z \rangle = \alpha \int_\Omega y \cdot z \, dx + \nu \int_\Omega \nabla y \cdot \nabla z \, dx + \int_\Omega y \cdot (\delta v \cdot \nabla) v \, dx + \int_\Omega y \cdot (z \cdot \nabla) v \, dx.
\] (3.20)

It has, therefore, a purely integral representation, a property which is of major importance in view of finite element implementations of the conjugate gradient algorithm to be described in the following section.

### 3.4. Conjugate gradient solution of the least squares problem (3.16)

#### 3.4.1. Description of the algorithm

**Step 0: Initialization**

\[
u^0 \in V_g, \quad \text{given};
\] (3.21)

we define then \( g^0, w^0 \in V_0 \) by
\[
\begin{aligned}
g^0 \in V_0, \\
\alpha \int_\Omega g^0 \cdot z \, dx + \nu \int_\Omega \nabla g^0 \cdot \nabla z \, dx = \langle J'(u^0), z \rangle, \quad \forall z \in V_0,
\end{aligned}
\] (3.22)

\[
w^0 = g^0,
\] (3.23)

respectively. \( \square \)

Then for \( m \geq 0 \), assuming that \( u^m, g^m, w^m \), are known, we obtain \( u^{m+1}, g^{m+1}, w^{m+1} \) as follows:
Step 1: Descent

\[
\left\{ \begin{array}{l}
\text{Find } \lambda_m \in \mathbb{R} \text{ such that } \\
J(\mathbf{u}^m - \lambda_m \mathbf{w}^m) \leq J(\mathbf{u}^m - \lambda \mathbf{w}^m), \quad \forall \lambda \in \mathbb{R}, \\
\mathbf{u}^{m+1} = \mathbf{u}^m - \lambda_m \mathbf{w}^m.
\end{array} \right.
\]  

(3.24)

Step 2: Calculation of the new descent direction

\[
\left\{ \begin{array}{l}
\text{Find } \mathbf{g}^{m+1} \in V_0 \text{ such that } \\
\alpha \int_{\Omega} \mathbf{g}^{m+1} \cdot \zeta \, d\mathbf{x} + \nu \int_{\Omega} \nabla \mathbf{g}^{m+1} \cdot \nabla \zeta \, d\mathbf{x} = \langle J'(\mathbf{u}^{m+1}), \zeta \rangle, \quad \forall \zeta \in V_0
\end{array} \right.
\]  

(3.26)

and compute either

\[
\gamma_m = \frac{\alpha \int_{\Omega} \mathbf{g}^{m+1} \cdot \mathbf{g}^{m+1} \, d\mathbf{x} + \nu \int_{\Omega} \nabla \mathbf{g}^{m+1} \cdot \nabla \mathbf{g}^{m+1} \, d\mathbf{x}}{\alpha \int_{\Omega} \mathbf{g}^m \cdot \mathbf{g}^m \, d\mathbf{x} + \nu \int_{\Omega} \nabla \mathbf{g}^m \cdot \nabla \mathbf{g}^m \, d\mathbf{x}}.
\]  

(3.27a)

or

\[
\gamma_m = \frac{\alpha \int_{\Omega} (\mathbf{g}^{m+1} - \mathbf{g}^m) \cdot \mathbf{g}^{m+1} \, d\mathbf{x} + \nu \int_{\Omega} \nabla (\mathbf{g}^{m+1} - \mathbf{g}^m) \cdot \nabla \mathbf{g}^{m+1} \, d\mathbf{x}}{\alpha \int_{\Omega} \mathbf{g}^m \cdot \mathbf{g}^m \, d\mathbf{x} + \nu \int_{\Omega} \nabla \mathbf{g}^m \cdot \nabla \mathbf{g}^m \, d\mathbf{x}}.
\]  

(3.27b)

or

\[
\gamma_m = \frac{\alpha \int_{\Omega} \mathbf{w}^m \cdot (\mathbf{g}^m - \mathbf{g}^{m+1}) \, d\mathbf{x} + \nu \int_{\Omega} \nabla \mathbf{w}^m \cdot \nabla (\mathbf{g}^m - \mathbf{g}^{m+1}) \, d\mathbf{x}}{\alpha \int_{\Omega} \mathbf{w}^m \cdot (\mathbf{g}^m - \mathbf{g}^{m+1}) \, d\mathbf{x} + \nu \int_{\Omega} \nabla \mathbf{w}^m \cdot \nabla (\mathbf{g}^m - \mathbf{g}^{m+1}) \, d\mathbf{x}}.
\]  

(3.27c)

and set

\[
\mathbf{w}^{m+1} = \mathbf{g}^{m+1} + \gamma_m \mathbf{w}^m.
\]  

(3.28)

Do \( m = m + 1 \) and go to (3.24).

Conjugate gradient algorithms can be quite effective as solution methods for very large scale nonlinear problems. Some basic references are refs. [25,26] (see also ref. [1 chap. 7] and refs. [27,28] for further references and applications). Algorithm (3.21)-(3.28) is known as a
Fletcher–Reeves (resp. Polak–Ribiere, resp. Beale) conjugate gradient algorithm if one uses (3.27a) (resp. (3.27b), resp (3.27c)) to compute $\gamma_m$. From numerical experiments it appears that Polak–Ribiere and Beale algorithms have, in general, better convergence properties than the Fletcher–Reeves algorithm (there are, however, counter examples).

As we shall see in section 3.4.2, applying algorithm (3.21)–(3.28) to solve the least squares problem (3.16) requires the solution at each iteration of exactly three Dirichlet systems (i.e. $3N$ scalar Dirichlet problems) associated to the elliptic operator $\alpha I - \nu \Delta$.

**Remark 3.1:** A most important step when making use of algorithm (3.21)–(3.28) to solve the least squares problem (3.16), is the calculation of $\langle J'(u_m^{m+1}), z \rangle$ at each iteration. From the calculation of $J'$, done in section 3.3 to obtain $\langle J'(u_m^{m+1}), z \rangle$, we should proceed as follows:

(i) We compute $\gamma_m^{m+1}$, associated to $u_m^{m+1}$ by (3.14)–(3.15), as indicated in section 3.4.2., below.
(ii) We obtain then $\langle J'(u_m^{m+1}), z \rangle$ by taking $v = u_m^{m+1}$ and $y = \gamma_m^{m+1}$ in (3.20).

### 3.4.2. Calculation of $\lambda_m$. Further comments on algorithm (3.21)–(3.28)

A problem of practical importance is the calculation of $\lambda_m$. Let's denote by $y_m(\lambda)$ the solution of (3.14), (3.15) associated to $v = u_m - \lambda w_m$. We clearly have

$$y_m^m(0) = y_m^m, \quad y_m^m(\lambda) = y_m^{m+1},$$

and also

$$y_m^m(\lambda) = y_m^m - \lambda y_1^m + \lambda^2 y_2^m,$$

where $y_1^m$, $y_2^m$ are the solutions of

$$\begin{cases}
\alpha y_1^m - \nu \Delta y_1^m = \alpha w_m - \nu \Delta w_m + (u_m \cdot \nabla)w_m + (w_m \cdot \nabla)u_m & \text{in } \Omega, \\
y_1^m = 0 & \text{on } \Gamma,
\end{cases}$$

and

$$\begin{cases}
\alpha y_2^m - \nu \Delta y_2^m = (w_m \cdot \nabla)w_m & \text{in } \Omega, \\
y_2^m = 0 & \text{on } \Gamma,
\end{cases}$$

respectively. Since

$$J(u_m - \lambda w_m) = \frac{1}{2} \int_{\Omega} \{ \alpha | y_m^m(\lambda) |^2 + \nu | \nabla y_m^m(\lambda) |^2 \} \, dx,$$

the function $\lambda \to J(u_m - \lambda w_m)$ is, from (3.30), a quartic polynomial in $\lambda$ that we shall denote by $J_m(\lambda)$ and $\lambda_m$ is, therefore, a solution of the cubic equation

$$J_m'(\lambda) = 0.$$

We shall use the standard Newton's method to compute $\lambda_m$ from (3.34), starting from $\lambda = 0$. The resulting algorithm is given by

$$\lambda^0 = 0.$$
Then for \( k \geq 0 \), we obtain \( \lambda^{k+1} \) from \( \lambda^k \) by

\[
\lambda^{k+1} = \lambda^k - j_m' (\lambda^k) / j_m'' (\lambda^k).
\] (3.36)

In our calculations we always observed a very fast convergence of algorithm (3.35), (3.36). Once \( \lambda_m \) is known, we know \( y^{m+1} \) since (from (3.29)) \( y^{m+1} = y^m (\lambda_m) \). If we count now the number of Dirichlet systems for \( \alpha I - \nu \Delta \) to be solved at each iteration, we observe that we have to solve only three such systems; namely (3.31), (3.32), and then (3.26) (to obtain \( g^{m+1} \)). This number is optimal for a nonlinear problem since the solution of a linear problem, by least squares–preconditioned conjugate gradient algorithms, requires the solution at each iteration of two linear systems associated to the preconditioning operator.

From the above remarks, it appears clearly that the practical implementation of algorithm (3.21)–(3.28) will require an efficient (direct or iterative) elliptic solver, like those discussed in refs. [29,30].

As a final comment, we would like to mention that algorithm (3.21)–(3.28) (in fact, its finite dimensional variants) is quite efficient when used in combination with the operator splitting methods of section 2 to solve the test problems of section 8. Three to five iterations suffice to reduce the value of the cost function by a factor of \( 10^4 \) to \( 10^6 \). However, in view of other applications, we have been testing some of these methods, such as the methods discussed in refs. [31,32], which combine the features of conjugate gradient and quasi-Newton algorithms. The results obtained for transonic potential flows have been quite impressive (see ref. [33]), but so far, quite disappointing for the Navier–Stokes equations (at least for the test cases we have been considering).

**Remark 3.2:** As stopping test for algorithm (3.21)–(3.28), we have been using

\[
J(u^m)/J(u^0) \leq \varepsilon,
\] (3.37)

where \( \varepsilon \) is a “small” positive number. Another test could have been

\[
\| g^m \| / \| g^0 \| \leq \varepsilon
\] (3.38)

with

\[
\| g^m \| = \left( \alpha \int_{\Omega} |g^m|^2 \, dx + \nu \int_{\Omega} |\nabla g^m|^2 \, dx \right)^{1/2}.
\]

We found (3.37) more interesting, however, since (3.38), unlike (3.37), cannot discriminate between global and local minima of \( J \).

4. Solution of the “quasi” Stokes linear subproblems

4.1. Generalities. Synopsis

At each full step of the splitting methods discussed in section 2.3, we have to solve one or two
linear problems of the following type

\[
\begin{aligned}
&\alpha u - \nu \Delta u + \nabla p = f \quad \text{in } \Omega, \\
&\nabla \cdot u = 0 \quad \text{in } \Omega, \\
&u = g \quad \text{on } \Gamma \quad \left(\text{with} \int_{\Omega} g \cdot n \, d\Gamma = 0\right),
\end{aligned}
\]

(4.1)

where \( \alpha \) and \( \nu \) are two positive parameters and where \( f \) and \( g \) are two given functions defined on \( \Omega \) and \( \Gamma \), respectively. We recall that if \( f \) and \( g \) are sufficiently smooth, then problem (4.1) has a unique solution in \( V_g \times (L^2(\Omega) / \mathbb{R}) \) (with \( V_g \) still defined by (3.5) and \( p \in L^2(\Omega) / \mathbb{R} \) means that \( p \) is defined only to within an arbitrary constant). We shall describe below several iterative methods for solving (4.1) which are quite easy to implement using finite element methods (more details are given in refs. [1,9,10,20], together with convergence proofs).

4.2. A first iterative method for solving (4.1)

This method is quite classical and is defined as follows

\[
p^0 \in L^2(\Omega), \quad \text{given;}
\]

(4.2)

then for \( m \geq 0 \), define \( u^m \) and \( p^{m+1} \) from \( p^m \) by

\[
\begin{aligned}
&\alpha u^m - \nu \Delta u^m = f - \nabla p^m \quad \text{in } \Omega, \\
u^m = g \quad \text{on } \Gamma,
\end{aligned}
\]

(4.3)

\[
p^{m+1} = p^m - \rho \nabla \cdot u^m.
\]

(4.4)

Concerning the convergence of algorithm (4.2)-(4.4), we have the following

**Proposition 4.1:** Suppose that

\[
0 < \rho < 2\nu/N;
\]

(4.5)

we have then

\[
\lim_{m \to +\infty} \{ u^m, p^m \} = \{ u, p_0 \} \quad \text{in } (H^1(\Omega))^N \times L^2(\Omega)
\]

(4.6)

where \( \{ u, p_0 \} \) is the solution of (4.1) such that

\[
\int_{\Omega} p_0 \, dx = \int_{\Omega} p^0 \, dx.
\]

(4.7)

Moreover, the convergence is linear (i.e. the sequence \( \| u^m - u \|_{(H^1(\Omega))^N} \) and \( \| p^m - p_0 \|_{L^2(\Omega)} \) converge to zero, as fast, at least, as geometric sequences).

See ref. [1] for a proof of proposition 4.1.
Remark 4.1: When using algorithm (4.2)-(4.4) to solve (4.1), we have to solve at each iteration, \( N \) uncoupled scalar Dirichlet problems for \( \alpha I - \nu \Delta \), to obtain \( u^m \) from \( p^m \). We see, again (as in section 3), the importance of having efficient Dirichlet solvers for \( \alpha I - \nu \Delta \).

Remark 4.2: Algorithm (4.2)-(4.4) is related to the so-called method of artificial compressibility of Chorin--Yancenko. Indeed, we can view (4.4) as obtained by a time discretization process from the equation

\[
\frac{\partial p}{\partial t} + \nabla \cdot u = 0
\]

(\( \rho \) being the size of the time discretization step).

4.3. Conjugate gradient methods for problem (4.2)

4.3.1. Generalities. Synopsis

The above algorithm (4.2)-(4.4) is, in fact, the simplest one of a whole family of descent methods, including among them conjugate gradient algorithms. To justify the use of conjugate gradient methods for solving problem (4.1), we shall show (in section 4.3.2) that the pressure \( p \) is a solution of a functional equation associated to a self-adjoint and strongly elliptic operator. Such an equation can be solved, therefore, by conjugate gradient methods like those described in sections 4.3.3, 4.3.4.

4.3.2. A functional equation satisfied by the pressure

We suppose \( \Omega \) is bounded in \( \mathbb{R}^N \). Let's define \( H \) by

\[
H = \left\{ q \mid q \in L^2(\Omega), \int_{\Omega} q(x) \, dx = 0 \right\},
\]

(4.8)

and then \( A : L^2(\Omega) \to L^2(\Omega) \) by:

Take \( q \in L^2(\Omega) \) and solve

\[
\begin{cases}
\alpha u_q - \nu \Delta u_q = -\nabla q & \text{in } \Omega, \\
u u_q = 0 & \text{on } \Gamma
\end{cases}
\]

(4.9)

in \( (H^1_0(\Omega))^N \); let then

\[
A q = \nabla \cdot u_q.
\]

(4.10)

Since

\[
\int_{\Omega} A q \, dx = \int_{\Omega} \nabla \cdot u_q \, dx = \int_{\Gamma} u_q \cdot n \, d\Gamma = 0,
\]

we have \( A q \in H, \forall q \in L^2(\Omega) \). We have also

\[
\int_{\Omega} (A q) q' \, dx = \alpha \int_{\Omega} u_q \cdot u_q' \, dx + \nu \int_{\Omega} \nabla u_q \cdot \nabla u_q' \, dx, \quad \forall q, \, q' \in L^2(\Omega),
\]

(4.11)

from which we can deduce the following theorem.
Theorem 4.1: The operator \( A \) is self-adjoint and is a strongly elliptic isomorphism from \( H \) onto \( H \).

See ref. [34] for a proof of this result. Define now \( u_0 \in V_g \) by

\[
\begin{align*}
\alpha u_0 - v \Delta u_0 &= f \quad \text{in} \quad \Omega, \\
u_0 &= g \quad \text{on} \quad \Gamma.
\end{align*}
\]

We clearly have, from the definition of \( A \) and from (4.1), (4.12), that

\[
Ap = -\nabla \cdot u_0.
\]

From the properties of \( A \), it is tempting to solve (4.13) (and therefore (4.1)) by conjugate gradient algorithms operating in the spaces \( L^2(\Omega) \) and \( H \). Such algorithms are discussed in sections 4.3.3, 4.3.4.

4.3.3. Conjugate gradient methods for a class of linear variational problems

Let \( V \) be a real Hilbert space whose scalar product and associated norm are denoted by

\[
(\cdot, \cdot) \quad \text{and} \quad \| \cdot \|,
\]

respectively. Consider now the following mathematical objects:

(i) \( a: V \times V \to \mathbb{R} \), a bilinear, continuous functional which is also \( V \)-elliptic (i.e., there exists \( \alpha > 0 \) such that \( a(v, v) \geq \alpha \| v \|^2, \forall v \in V \)),

(ii) \( L: V \to \mathbb{R} \), linear and continuous.

To \( a(\cdot, \cdot) \) and \( L(\cdot) \), we associate the following (linear variational) problem

\[
\begin{align*}
\text{Find} \quad u \in V \quad \text{such that} \\
a(u, v) &= L(v), \quad \forall v \in V.
\end{align*}
\]

(4.14)

It is a classical result that problems such as (4.14) have a unique solution if the above hypotheses on \( a(\cdot, \cdot) \) and \( L(\cdot) \) holds (see, for example, [1, appendix 1] for a proof).

Suppose now that \( a(\cdot, \cdot) \) is symmetric, i.e.

\[
a(v, w) = a(w, v), \quad \forall v, w \in V.
\]

(4.15a)

In such a case, it is easily shown that problem (4.14) is also equivalent to the following minimization problem

\[
\begin{align*}
\text{Find} \quad u \in V, \\
j(u) \leq j(v), \quad \forall v \in V,
\end{align*}
\]

(4.15a)

with

\[
j(v) = \frac{1}{2} a(v, v) - L(v).
\]

(4.15b)
In the symmetric case, the above equivalence result suggests solving (4.14), via (4.13), by a conjugate gradient method, like, for example, the one described below.

**Description of a conjugate gradient method for solving (4.14)**

**Step 0: Initialization**

\[ u^0 \in V \text{ is given;} \]  

solve then

\[
\begin{cases}
  g^0 \in V, \\
  (g^0, v) = a(u^0, v) - L(v), \quad \forall v \in V;
\end{cases}
\]

if \( g^0 = 0 \) (or if \( \| g^0 \| \) is sufficiently small) take \( u = u^0 \); if not, set

\[ w^0 = g^0. \]

Then for \( m \geq 0 \) assuming that \( u^m, g^m, w^m \) are known, compute \( u^{m+1}, g^{m+1}, w^{m+1} \) as follows:

**Step 1: Descent**

Compute

\[
\rho_m = \frac{(g^m, w^m)}{a(w^m, w^m)} = \frac{\|g^m\|^2}{a(w^m, w^m)},
\]

and take

\[ u^{m+1} = u^m - \rho_m w^m. \]

**Step 2: Test of the convergence and calculation of the new descent direction**

Solve

\[
\begin{cases}
  g^{m+1} \in V, \\
  (g^{m+1}, v) = (g^m, v) - \rho_m a(w^m, v), \quad \forall v \in V;
\end{cases}
\]

if \( g^{m+1} = 0 \), or if \( \| g^{m+1} \| \) is sufficiently small, take \( u = u^{m+1} \); if not, compute

\[ \gamma_m = \|g^{m+1}\|^2 / \|g^m\|^2 \]

and define

\[ w^{m+1} = g^{m+1} = \gamma_m w^m. \]

Do \( m = m + 1 \) and go to (4.19).
Proving the convergence of algorithm (4.16)–(4.23) is not too difficult (see, e.g. ref. [26] for such a proof). What is more interesting, indeed, is that one has the following estimate for the speed of convergence (this is also proved in ref. [26]):

$$\| u^m - u \| \leq c \left( \frac{\nu(a) - 1}{\nu(a) + 1} \right)^m,$$

(4.24)

where $c$ depends on $u^0$. In (4.24), $\nu(a)$ denotes the condition number of $a(\cdot, \cdot)$, namely

$$\nu(a) = \| A \| \| A^{-1} \|,$$

(4.25)

where $A$ is the isomorphism from $V$ onto $V$ uniquely defined by

$$a(v, w) = (Av, w), \quad \forall v, w \in V.$$

(4.26)

Despite its apparent simplicity, algorithm (4.16)–(4.23) is one of the most useful tools in Scientific Computing. The number of situations to which it can be applied is enormous and would fill a very large book by itself (see, e.g. refs. [1,35] for some applications). In section 4.3.4 just below, we shall apply algorithm (4.16)–(4.23) to the solution of problem (4.1), via the solution of the linear eq. (4.13).

4.3.4. Application of the conjugate gradient algorithm (4.16)–(4.23) to the solution of the Stokes problem (4.1), via the solution of (4.13)

Let us denote by $p$ the pressure in (4.1) belonging to

$$H = \left\{ q | q \in L^2(\Omega), \quad \int_{\Omega} q \, dx = 0 \right\}.$$

It follows from section 4.3.2 that $p$ satisfies (4.13), or equivalently:

$$\begin{cases} p \in H, \\ \int_{\Omega} (Ap) q \, dx = - \int_{\Omega} \nabla \cdot u_0 q \, dx, \quad \forall q \in H. \end{cases}$$

(4.27)

From the properties of $A$ discussed in section 4.3.2 and particularly theorem 4.1, the bilinear form

$$\{ q, q' \} \to \int_{\Omega} (Aq) q' \, dx$$

is continuous, symmetric and $H$-elliptic over $H \times H$. Since the linear functional

$$q \to \int_{\Omega} \nabla \cdot u_0 q \, dx$$
is clearly continuous over $H$, problem (4.27) is a particular case of problem (4.14) and therefore, can be solved by the conjugate gradient algorithm (4.16)-(4.23). Before describing the resulting algorithm, we have to decide on the scalar product to be used over $H$. In that direction, let’s consider first two real parameters $a$ and $b$, such that

$$a > 0 \quad \text{and} \quad b > 0.$$ 

We define now the operator $B$ from $H$ into $H$ by

$$Bq = aq + b\phi_q,$$

where $\phi_q$ is the unique solution in $H \cap H^1(\Omega)$ of the Neumann problem

$$\begin{aligned}
-\Delta \phi_q &= q \quad \text{in} \quad \Omega, \\
\partial \phi_q / \partial n &= 0 \quad \text{on} \quad \Gamma;
\end{aligned} \quad (4.29)$$

(since $\int_\Omega q \, dx = 0$, the Neumann problem (4.29) has an infinity of solutions in $H^1(\Omega)$, all defined to within an arbitrary additive constant and therefore, there is a unique solution of (4.29) such that $\int_\Omega \phi_q \, dx = 0$). Operator $B$ is clearly continuous from $H$ into $H$ ($H$ being equipped with the classical $L^2(\Omega)$-norm

$$\|q\|_{L^2(\Omega)} = \left( \int_\Omega |q(x)|^2 \, dx \right)^{1/2}.$$ 

In fact, operator $B$ is a self-adjoint strongly elliptic isomorphism from $H$ onto $H$. To prove these properties, it suffices to prove that the continuous bilinear form

$$d: H \times H \to \mathbb{R}$$

defined by

$$d(q, q') = \int_\Omega (Bq) q' \, dx$$ 

(4.30)

is symmetric and $H$-elliptic *. We have, with obvious notation,

$$d(q, q') = a \int_\Omega qq' \, dx + b \int_\Omega \phi_q q' \, dx = a \int_\Omega qq' \, dx - b \int_\Omega \phi_q \Delta \phi_q \, dx$$

$$= a \int_\Omega qq' \, dx + b \int_\Omega \nabla \phi_q \cdot \nabla \phi_q' \, dx, \quad \forall q, q' \in H.$$ 

(4.31)

The symmetry of $d(\cdot, \cdot)$ is obvious from (4.31). We also have

$$d(q, q) = a \int_\Omega |q|^2 \, dx + b \int_\Omega |\nabla \phi_q|^2 \, dx \geq a \int_\Omega |q|^2 \, dx, \quad \forall q \in H.$$ 

(4.32)

The $H$-ellipticity of $d(\cdot, \cdot)$ is clear from (4.32).

* I.e. there exists $\beta > 0$ such that $d(q, q) \geq \beta \|q\|_{L^2(\Omega)}^2, \forall q \in H.$
Let’s denote by $S$ the inverse of operator $B$, i.e.

$$S = B^{-1}.$$  \hfill (4.33)

From the properties of $B$, operator $S$ is also a strongly elliptic and self-adjoint isomorphism from $H$ onto $H$, implying that the bilinear form

$$\{q, q'\} \rightarrow \int_\Omega (Sq) q' \, dx$$

defines over $H$, a scalar product such that the associated norm is equivalent to the classical $L^2(\Omega)$-norm. From now on, we shall use the Hilbert structure over $H$ associated with the scalar product $(\cdot , \cdot)_H$ defined by

$$\langle q, q' \rangle_H = \int_\Omega (Sq) q' \, dx = \langle Sq, q' \rangle_{L^2(\Omega)}.$$ \hfill (4.34)

Indeed, we shall use $(\cdot , \cdot)_H$ for the solution of problem (4.1) via the conjugate gradient solution of problem (4.13) by algorithm (4.16)-(4.23).

The above metric was introduced, using a different presentation, in refs. [36,37] and is discussed in further detail in ref. [38] (where more general boundary conditions are considered).

**Description of the conjugate gradient algorithm**

We follow the presentation in ref. [38, chap. 2]. Algorithm (4.16)-(4.23), applied to the solution of the Stokes problem (4.1), yields, if one uses the scalar product defined by (4.34), the following algorithm:

**Step 0: Initialization**

$$p^0 \in H, \text{ given;}$$ \hfill (4.35)

solve then the elliptic system

$$\begin{cases}
\alpha u^0 - \nu \Delta u^0 = f - \nabla p^0 \text{ in } \Omega, \\
u u^0 = g \text{ on } \Gamma.
\end{cases} \hfill (4.36)$$

If $\nabla \cdot u^0 = 0$ (or, in practice, is sufficiently small), then take $u = u^0$ and $p = p^0$; if this is not the case, compute the solution of the Neumann’s problem

$$\begin{cases}
-\Delta \psi^0 = \nabla \cdot u^0 \text{ in } \Omega, \\
\psi^0 / \partial n = 0 \text{ on } \Gamma
\end{cases} \hfill (4.37)$$
which satisfies $\int_\Omega \psi^0 \, dx = 0$, and set
\[
g^0 = a \nabla \cdot u^0 + b \psi^0, \quad (4.38)
\]
\[
w^0 = g^0. \quad (4.39)
\]

Then for $m \geq 0$, assuming that $p^m, u^m, g^m, w^m$ are known, compute $p^{m+1}, u^{m+1}, g^{m+1}, w^{m+1}$ as follows:

**Step 1: Descent**

Solve the elliptic system
\[
\begin{aligned}
\alpha \chi^m - \nu \Delta \chi^m &= - \nabla w^m \quad \text{in } \Omega, \\
\chi^m &= 0 \quad \text{on } \Gamma,
\end{aligned}
\]
and compute
\[
\rho_m = \frac{\int_\Omega \nabla \cdot u^m w^m \, dx}{\int_\Omega \nabla \cdot \chi^m w^m \, dx} = \frac{\int_\Omega \nabla \cdot u^m g^m \, dx}{\int_\Omega \nabla \cdot \chi^m w^m \, dx},
\]
\[
p^{m+1} = p^m - \rho_m w^m, \quad (4.42)
\]
\[
u^{m+1} = u^m - \rho_m \chi^m. \quad (4.43)
\]

**Step 2: Testing the convergence and calculation of the new descent direction**

If $\nabla \cdot u^{m+1} = 0$ (or, in practice, is sufficiently small), take $u = u^{m+1}, \ p = p^{m+1}$; if this is not the case, solve the Neumann's problem
\[
\begin{aligned}
\Delta \phi^m &= \nabla \cdot \chi^m \quad \text{in } \Omega, \\
\partial \phi^m / \partial n &= 0 \quad \text{on } \Gamma
\end{aligned}
\]
with $\int_\Omega \phi^m \, dx = 0$, and take
\[
g^{m+1} = g^m - \rho_m (a \nabla \cdot \chi^m + b \phi^m),
\]
\[
\gamma_m = \frac{\int_\Omega g^{m+1} \nabla \cdot u^{m+1} \, dx}{\int_\Omega g^m \nabla \cdot u^m \, dx},
\]
\[
w^{m+1} = g^{m+1} + \gamma_m w^m, \quad (4.47)
\]

Do $m = m + 1$ and go to (4.41).
If follows from Fourier Analysis (cf. refs. [37,38] for details) that the optimal choice for \(a\) and \(b\) is provided by

\[
a = v, \quad b = \alpha.
\]  

(4.48)

With such choices, the numerical experiments reported in refs. [36–38] show that the convergence of algorithm (4.35)–(4.47) is very fast (3 to 4 iterations at most) and remarkably independent of the Reynolds number (as soon as \(\alpha \sim 1/\Delta t\) is sufficiently large).

A generalization of algorithm (4.35)–(4.47), which is able to handle boundary conditions more complicated than Dirichlet's, is discussed in ref. [38].

We observe that algorithm (4.35)–(4.47) requires the solution, at each iteration, of only one Dirichlet system, namely (4.40). This algorithm is therefore, almost as economical per iteration as algorithm (4.2)–(4.4), but since it requires fewer iterations, we should use (4.35)–(4.47) instead of (4.2)–(4.4).

4.4. An exact penalty method for solving the Stokes problem (4.1)

This method is, in fact, a generalization of algorithm (4.2)–(4.4) and is related to the augmented Lagrangian methods discussed in the monograph of Glowinski and Le Tallec [39] (see also Fortin and Glowinski [40], Glowinski [1, chap. 7]). It is defined as follows (with \(r\) a positive parameter):

\[
p^0 \in L^2(\Omega), \quad \text{given};
\]  

(4.49)

then for \(m \geq 0\), define \(u^m\) and \(p^{m+1}\) from \(p^m\) by

\[
\begin{align*}
\alpha u^m - v \Delta u^m - r \nabla (\nabla \cdot u^m) &= f - \nabla p^m \\
u^m &= g \quad \text{on } \Gamma, \\
p^{m+1} &= p^m - \rho \nabla \cdot u^m.
\end{align*}
\]  

(4.50)

(4.51)

Concerning the convergence of algorithm (4.49)–(4.51), we should prove (see, e.g. ref. [1, chap. 7] for such a proof) the following:

Proposition 4.2: Suppose that

\[
0 < \rho < 2(r + v/N);
\]  

(4.52)

then the convergence result (4.6) holds for \(\{u^m, p^m\}\) the convergence being still linear. Moreover, if one uses \(\rho = r\) we have

\[
\|p^m - p_0\|_{L^2(\Omega)} \leq \|p^0 - p_0\|_{L^2(\Omega)} \left( \frac{\|A^{-1}\|}{r} \right)^m, \quad \forall m \geq 0,
\]

where \(A\) is the operator defined by (4.9), (4.10); a similar estimate holds for \(\|u^m - u\|_{(H^1(\Omega))^N}\) if \(\rho = r\).
Remark 4.3. (About the choice of $\rho$ and $r$): We should use $\rho = r$ in practice, since (from the above proposition) the convergence ratio of algorithm (4.49)–(4.51) is $O(r^{-1})$ for large values of $r$. In many applications, taking $r = 10^4\rho$, we have a practical convergence of algorithm (4.49)–(4.51) in 3 to 4 iterations. There is, however, a practical upper bound for $r$. This follows from the fact that for too large values of $r$, problem (4.50) will be ill-conditioned and its practical solution sensitive to round-off errors.

Remark 4.4: Problem (4.50) is more complicated to solve in practice than problem (4.3) (or (4.40)), since the components of $u''$ are coupled by the linear term $\nabla(\nabla \cdot u'')$. Actually, the elliptic partial differential operator in the left hand side of (4.50) is very close to the linear elasticity operator and close variants of it occur naturally in compressible (see part II) and/or turbulent viscous flow problems.

Remark 4.5: Other methods for solving the “quasi” Stokes problem (4.1) are discussed in refs. [1,10,20,41,42].

5. Finite element approximation of the time dependent Navier–Stokes equations

We shall describe in this section a specific class of finite element approximations for the time dependent Navier–Stokes equations. Actually, these methods, which lead to continuous approximations for both pressure and velocity, are fairly simple and some of them have been known for years. They have been advocated, for example, by Hood and Taylor (see ref. [43]). Other finite element approximations of the incompressible Navier–Stokes equations can be found in refs. [1,6,9,10,44] (see also the references therein).

The two fundamental references (in our opinion) concerning the convergence of the fully discrete Navier–Stokes equations (finite elements in space, finite differences in time) are refs. [45,46].

5.1. Basic hypotheses. Fundamental discrete spaces

We suppose that $\Omega$ is a bounded polygonal domain of $R^2$. With $T_h$ a standard finite element triangulation of $\Omega$, and $h$ the maximal length of the edges of the triangles of $T_h$, we introduce the following discrete spaces (with $P_k = $ space of the polynomials in two variables of degree $\leq k$):

$$H^1_h = \{ q_h \mid q_h \in C^0(\bar{\Omega}), \quad q_h \mid T \in P_1, \quad \forall T \in T_h \},$$

$$V_h = \{ v_h \mid v_h \in C^0(\bar{\Omega}) \times C^0(\bar{\Omega}), \quad v_h \mid T \in P_2 \times P_2, \quad \forall T \in T_h \},$$

$$V_{0h} = \{ v_h \mid v_h \in V_h, \quad v_h = 0 \quad \text{on} \quad \Gamma \} = V_h \cap (H^1_0(\Omega))^2.$$

Two useful variants of $V_h$ (and $V_{0h}$) are obtained as follows: either

$$V_h = \{ v_h \mid v_h \in C^0(\bar{\Omega}) \times C^0(\bar{\Omega}), \quad v_h \mid T \in P_1 \times P_1, \quad \forall T \in T_h \}.$$
In (5.4), $\hat{T}_h$ is the triangulation of $\Omega$ obtained from $T_h$ by joining the midpoints of the edges of $T \in T_h$, as shown in fig. 5.1. We have the same global number of unknowns if we use $V_h$ defined by either (5.2) or (5.4), however, the matrices encountered in the second case are more compact and sparse.

In (5.5), $P_{1T}^*$ is the subspace of $P_3$ defined as follows

$$
P_{1T}^* = \{ q \mid q = q_1 + \lambda \phi_T, \quad \text{with} \quad q_1 \in P_1, \lambda \in \mathbb{R}, \quad \text{and} \quad \phi_T \in P_3, \quad \phi_T = 0 \quad \text{on} \quad \partial T, \quad \phi_T(G_T) = 1 \},$$

where, in (5.6), $G_T$ is the centroid of $T$ (see fig. 5.2 above). A function like $\phi_T$ is usually called a bubble-function.

5.2. Approximation of the boundary conditions

If the boundary conditions are defined by

$$u = g \quad \text{on} \quad \Gamma, \quad \int_{\Gamma} g \cdot n \, d\Gamma = 0,$$

it is of fundamental importance to approximate $g$ by $g_h$ such that

$$\int_{\Gamma} g_h \cdot n \, d\Gamma = 0.$$

Let's discuss the construction of such a $g_h$ (we follow here ref. [1, appendix 3]). For simplicity, we shall suppose that $g$ is continuous over $\Gamma$. We now define the space $\gamma V_h$ as

$$\gamma V_h = \{ \mu_h \mid \mu_h = \nu_h \mid \Gamma, \nu_h \in V_h \},$$

i.e., $\gamma V_h$ is the space of the traces on $\Gamma$ of those functions $\nu_h$ belonging to $V_h$. Actually, if $V_h$ is defined by (5.2), $\gamma V_h$ is also the space of those functions defined over $\Gamma$, taking their values in $\mathbb{R}^2$, continuous over $\Gamma$ and piecewise quadratic over the edges of $T_h$ contained in $\Gamma$. 

or (this space has been introduced in ref. [47])

$$V_h = \{ \nu_h \mid \nu_h \in C^0(\bar{\Omega}) \times C^0(\bar{\Omega}), \quad \nu_h \mid T \in P_{1T}^* \times P_{1T}^*, \quad \forall T \in T_h \}. \quad (5.5)$$
Our problem is to construct an approximation $g_h$ of $g$ such that

$$g_h \in \gamma V_h, \quad \int_{\Gamma} g_h \cdot n \, d\Gamma = 0. \quad (5.10)$$

If $\Pi_h g$ is the unique element of $\gamma V_h$, obtained by piecewise quadratic interpolation of $g$ over $\Gamma$, i.e., obtained from the values taken by $g$ at those nodes of $T_h$ belonging to $\Gamma$, we usually have $\int_{\Gamma} \Pi_h g \cdot n \, d\Gamma \neq 0$. To overcome this difficulty, we may proceed as follows:

(i) We define an approximation $n_h$ of $n$ as the solution of the following linear variational problem in $\gamma V_h$:

$$n_h \in \gamma V_h, \quad \int_{\Gamma} n_h \cdot \mu_h \, d\Gamma = \int_{\Gamma} n \cdot \mu_h \, d\Gamma, \quad \forall \mu_h \in \gamma V_h. \quad (5.11)$$

Problem (5.11) is, in fact, equivalent to a linear system whose matrix is sparse, positive definite, well-conditioned and quite easy to compute.

(ii) Then define $g_h$ by

$$g_h = \Pi_h g - \left( \int_{\Gamma} \Pi_h g \cdot n \, d\Gamma / \int_{\Gamma} n \cdot n_h \, d\Gamma \right) n_h. \quad (5.12)$$

It is easy to check that (5.11), (5.12) imply (5.10).

5.3. Space approximation of the time dependent Navier–Stokes equations

Using the spaces $H^1_h$, $V_h$ and $V_{0h}$, we approximate the time dependent Navier–Stokes equations as follows:

Find $\{ u_h(t), p_h(t) \} \in V_h \times H^1_h, \quad \forall t \geq 0,$ such that

$$\int_{\Omega} \frac{\partial u_h}{\partial t} \cdot v_h \, dx + \nu \int_{\Omega} \nabla u_h \cdot \nabla v_h \, dx + \int_{\Omega} (u_h \cdot \nabla) u_h \cdot v_h \, dx + \int_{\Omega} \nabla p_h \cdot v_h \, dx$$

$$= \int_{\Omega} f_h \cdot v_h \, dx, \quad \forall v_h \in V_{0h}, \quad (5.13)$$

$$\int_{\Omega} \nabla \cdot u_h q_h \, dx = 0, \quad \forall q_h \in H^1_h, \quad (5.14)$$

$$u_h = g_h \quad \text{on} \quad \Gamma, \quad (5.15)$$

$$u_h(x, 0) = u_{0h} \quad \text{(with} \quad u_{0h} \in V_h). \quad (5.16)$$

In (5.13)–(5.16), $f_h$, $u_{0h}$ and $g_h$ are convenient approximations of $f$, $u_0$ and $g$, respectively.
5.4. Time discretization by operator splitting methods

We consider now a fully discrete version of scheme (2.31)–(2.33) discussed in section 2.3.1. It is defined as follows (with $\Delta t$ as in section 2.3.1):

$$u_h^0 = u_{0h};$$

then for $n \geq 0$, compute (from $u_h^n$) $\{u_h^{n+1/2}, p_h^{n+1/2}\} \in V_h \times H_h^1$, and then $u_h^{n+1} \in V_h$, by solving

$$\left\{ \begin{array}{l}
\int_\Omega \frac{u_{h+1/2} - u_h^n}{\Delta t/2} \cdot v_h \, dx + \frac{\nu}{2} \int_\Omega \nabla u_{h+1/2} \cdot \nabla v_h \, dx + \int_\Omega \nabla p_{h+1/2} \cdot v_h \, dx \\
= \int_\Omega f_{h+1/2} \cdot v_h \, dx - \frac{\nu}{2} \int_\Omega \nabla u_h^n \cdot \nabla v_h \, dx - \int_\Omega (u_h^n \cdot \nabla) u_h^n \cdot v_h \, dx,
\end{array} \right. \quad (5.18a)$$

$$\forall v_h \in V_{0h},$$

$$\int_\Omega \nabla \cdot u_h^{n+1/2} q_h \, dx = 0, \quad \forall q_h \in H_h^1, \quad (5.18b)$$

$$u_h^{n+1/2} \in V_h, \quad p_h^{n+1/2} \in H_h^1, \quad u_h^{n+1/2} = g_h^{n+1/2} \quad \text{on } \Gamma, \quad (5.18c)$$

and then

$$\left\{ \begin{array}{l}
\int_\Omega \frac{u_{h+1} - u_h^{n+1/2}}{\Delta t/2} \cdot v_h \, dx + \frac{\nu}{2} \int_\Omega \nabla u_h^{n+1/2} \cdot \nabla v_h \, dx + \int_\Omega (u_h^{n+1/2} \cdot \nabla) u_h^{n+1} \cdot v_h \, dx \\
= \int_\Omega f_{h+1} \cdot v_h \, dx - \frac{\nu}{2} \int_\Omega \nabla u_h^{n+1/2} \cdot \nabla v_h \, dx - \int_\Omega \nabla p_h^{n+1/2} \cdot v_h \, dx, \quad \forall v_h \in V_{0h},
\end{array} \right. \quad (5.19a)$$

$$u_h^{n+1} \in V_h, \quad u_h^{n+1} = g_h^{n+1} \quad \text{on } \Gamma, \quad (5.19b)$$

respectively. The same techniques apply to the space discretization of scheme (2.34)–(2.37).

The solution of the subproblems encountered at each step of (5.17)–(5.19) can be obtained by discrete variants of the methods discussed in sections 3 and 4; see ref. [1] (and also ref. [48]) for more details about this part of the solution process.


6.1. Generalities. Synopsis

These last years have seen a significant decrease in the popularity of the solution methods for the Navier–Stokes equations using the stream function–vorticity formulation. We see two main
reasons for this phenomenon:
(i) These methods are really convenient for two-dimensional flows. The generalization to
three-dimensional flows, although possible, leads to complicated formulations.
(ii) The treatment of the boundary conditions is more delicate than with the velocity–pressure
formulation.
We have decided, however, to dedicate a full section to the stream function–vorticity
formulation of the Navier–Stokes equations since an abundant amount of literature exists on this
approach and because the difficulties associated with the boundary conditions can be largely
overcome nowadays (see, e.g. ref. [49] for the numerical treatment of the boundary conditions in
the stream function–vorticity formulation of the Stokes problem).
In section 6.2, we shall derive the stream function–vorticity formulation from the
velocity–pressure model. In section 6.3, we shall give a variational formulation of the new
equations, most useful in view of the time and space discretizations which are to be discussed in
sections 6.4 and 6.5, respectively. Some numerical results obtained by using the stream
function–vorticity formulation will be presented in section 8.

6.2. The stream function–vorticity formulation of the Navier–Stokes equations for incompressible
viscous flows

We use here the notation of section 1. We denote by $\Omega$ the two-dimensional flow region and
denote by $\Gamma$ its boundary. We suppose that $\Omega$ is $q$-connected with $q$ a nonnegative integer. The
possible holes (corresponding to obstacles to the flow) are denoted by $\omega_k$, $k = 1, \ldots, q$. If we
denote by $\Gamma$ the boundary of $\Omega$, we have (with the notation of fig. 6.1, where $q = 3$)

$$
\Gamma = \bigcup_{k=0}^{q} \Gamma_k.
$$

From section 1, unsteady flows in $\Omega$ of Newtonian, incompressible viscous fluids are modeled by
the following Navier–Stokes equations (written here in the velocity–pressure formulation)

$$
\frac{\partial u}{\partial t} - \nu \Delta u + (u \cdot \nabla) u + \nabla p = f \quad \text{in} \quad \Omega, \quad (6.1)
$$

$$
\nabla \cdot u = 0 \quad \text{in} \quad \Omega, \quad (6.2)
$$

Fig. 6.1.
with the initial condition
\[ u(x, 0) = u_0(x) \quad (\text{with } \nabla \cdot u_0 = 0). \] (6.3)

Concerning the boundary conditions, we shall consider only Dirichlet boundary conditions for the velocity, i.e.,
\[ u = g \quad \text{on } \Gamma. \] (6.4)

From the incompressibility property, \( g \) has to satisfy
\[ \int_{\Gamma} g \cdot n \, d\Gamma = 0. \]

In fact, we shall suppose that, for simplicity,
\[ \int_{\Gamma_k} g \cdot n \, d\Gamma = 0, \quad \forall k = 0, 1, \ldots, q. \] (6.5)

In order to derive a stream function-vorticity formulation from (6.1)-(6.4), we observe first that the incompressibility property \( \nabla \cdot u = 0 \) implies the existence of a (non-unique) vector \( \psi = \{0, 0, \psi\} \) such that
\[ u = \nabla \times \psi, \]
i.e.,
\[ u_1 = \frac{\partial \psi}{\partial x_2}, \quad u_2 = -\frac{\partial \psi}{\partial x_1}. \] (6.6)

A function such as \( \psi \) is called a stream function since the (obvious) relation
\[ u \cdot \nabla \psi = 0 \]
implies that \( \psi \) is constant along the streamlines.

If we define now \( \omega = \{0, 0, \omega\} \) by
\[ \omega = \nabla \times u, \] (6.7)
we clearly have
\[ \omega = \frac{\partial u_2}{\partial x_1} - \frac{\partial u_1}{\partial x_2}, \] (6.8)
which implies in turn (combined with (6.6))
\[ -\Delta \psi = \omega. \] (6.9)
The Poisson equation relating the stream function $\psi$ to the vorticity $\omega$ will be completed by the time dependent equation describing the evolution of $\omega$. To obtain such an equation, let’s take the curl of the momentum equation (6.1). We obtain then (with $f = \{f_1, f_2\}$):

$$\frac{\partial \omega}{\partial t} - \nu \Delta \omega + \frac{\partial \psi}{\partial x_2} \frac{\partial \omega}{\partial x_1} - \frac{\partial \psi}{\partial x_1} \frac{\partial \omega}{\partial x_2} = \frac{\partial f_2}{\partial x_1} - \frac{\partial f_1}{\partial x_2}$$

(6.10)

to which we associate the initial condition

$$\omega(x, 0) = \omega_0(x)$$

(6.11)

where $\{0, 0, \omega_0\} = \nabla \times u_0$.

Let’s derive now the boundary conditions associated with $\omega$ and $\psi$. At this stage, it is convenient to denote by $n_1$, $n_2$ the two coordinates of $n$ and to introduce

$$\tau = \{-n_2, n_1\},$$

which is a vector of unit length, tangent to $\Gamma$. We have then

$$\frac{\partial \psi}{\partial n} = \frac{\partial \psi}{\partial x_1} n_1 + \frac{\partial \psi}{\partial x_2} n_2 = u_1 n_2 - u_2 n_1$$

$$= -(u_1 \tau_1 + u_2 \tau_2) = -g \cdot \tau \quad \text{on} \quad \Gamma.$$

Similarly, we should prove that

$$\frac{\partial \psi}{\partial \tau} = \frac{\partial \psi}{\partial x_1} \tau_1 + \frac{\partial \psi}{\partial x_2} \tau_2 = u_1 n_1 + u_2 n_2$$

$$= u \cdot n = g \cdot n \quad \text{on} \quad \Gamma.$$

Thus $\psi$ satisfies the two following boundary conditions:

$$\frac{\partial \psi}{\partial n} = -g \cdot \tau \quad \text{on} \quad \Gamma,$$

(6.12)

$$\frac{\partial \psi}{\partial \tau} = g \cdot n \quad \text{on} \quad \Gamma.$$  

(6.13)

We observe that there is no simple boundary condition associated with $\omega$.

To summarize, the pair $\{\omega, \psi\}$ satisfies the following relations:

$$\frac{\partial \omega}{\partial t} - \nu \Delta \omega + \frac{\partial \psi}{\partial x_2} \frac{\partial \omega}{\partial x_1} - \frac{\partial \psi}{\partial x_1} \frac{\partial \omega}{\partial x_2} = \frac{\partial f_2}{\partial x_1} - \frac{\partial f_1}{\partial x_2} \quad \text{in} \quad \Omega,$$

(6.14a)

$$-\Delta \psi = \omega,$$  

(6.14b)

$$\omega(x, 0) = \omega_0(x),$$  

(6.14c)
We observe that
\[
\frac{\partial \psi}{\partial n} = -g \cdot \tau \quad \text{on} \quad \Gamma, \quad (6.14d)
\]
\[
\frac{\partial \psi}{\partial \tau} = g \cdot n \quad \text{on} \quad \Gamma. \quad (6.14e)
\]

In the next section, we shall discuss a variational formulation of (6.14) which is well suited to
time and space discretization.

6.3. Variational formulation of problem (6.14)

Multiplying (6.14a) (respectively (6.14b)) by \( \theta \in \Pi_0^1(\Omega) \) (respectively \( \phi \in H^1(\Omega) \)) and applying
Green's formula, we obtain
\[
\int_\Omega \frac{\partial \omega}{\partial t} \phi \, dx + \nu \int_\Omega \nabla \omega \cdot \nabla \phi \, dx + \int_\Omega \left( \frac{\partial \psi}{\partial x_2} \frac{\partial \omega}{\partial x_1} - \frac{\partial \psi}{\partial x_1} \frac{\partial \omega}{\partial x_2} \right) \phi \, dx
\]
\[\quad - \int_\Omega \left( f_1 \frac{\partial \theta}{\partial x_2} - f_2 \frac{\partial \theta}{\partial x_1} \right) \phi \, dx, \quad \forall \phi \in H^1_0(\Omega), \quad (6.15a)\]
\[
\omega (x, 0) = \omega_0 (x), \quad (6.15b)
\]
\[
\int_\Omega \nabla \psi \cdot \nabla \phi \, dx = \int_\Omega \omega \phi \, dx - \int_\Gamma g \cdot \tau \phi \, d\Gamma, \quad \forall \phi \in H^1(\Omega), \quad (6.15c)
\]
\[
\frac{\partial \psi}{\partial \tau} = g \cdot n \quad \text{on} \quad \Gamma. \quad (6.15d)
\]

which is equivalent to (6.14).

6.4. Time discretization of problem (6.14), (6.15)

6.4.1. Generalities. Synopsis

We shall discuss in this section the time discretization of problem (6.14), (6.15) by a
semi-implicit scheme, which is quite easy to implement in combination with finite element
approximations for the space discretization. Fully implicit schemes are described in ref. [50].

Before describing the time discretization of problem (6.14), (6.15), it is convenient to
"localize" \( \psi \). To achieve such a goal, we define \( \gamma_k, k = 0, \ldots, q \), by
\[
\gamma_k (m) = \int_{\Gamma_k} g \cdot n \, d\Gamma, \quad \forall m \in \Gamma_k. \quad (6.16)
\]
where $M_k$ is an arbitrary point of $\Gamma_k$. The function $\gamma_k$ vanishes at $M_k$ and is, from (6.5), continuous on $\Gamma_k$. Since $\psi$ is defined within an arbitrary constant, we require
\[
\psi = \gamma_0 \quad \text{on} \quad \Gamma_0
\] (6.17)
to fix this constant.

We introduce now, the two spaces $W_0$ and $W_g$ defined as follows:
\[
W_0 = \{ \phi | \phi \in H^1(\Omega), \quad \phi = 0 \quad \text{on} \quad \Gamma_0, \quad \partial \phi / \partial n = 0 \quad \text{on} \quad \Gamma_k, \quad \forall k = 1, \ldots, q \},
\] (6.18)
\[
W_g = \{ \phi | \phi \in H^1(\Omega), \quad \phi = \gamma_0 \quad \text{on} \quad \Gamma_0, \quad \partial \phi / \partial n = g \cdot n \quad \text{on} \quad \Gamma_k, \quad \forall k = 1, \ldots, q \}.
\] (6.19)

Both spaces $W_0$ and $W_g$ are closed in $H^1(\Omega)$.

Alternative definitions of $W_0$ and $W_g$ are given by
\[
W_0 = \{ \phi | \phi \in H^1(\Omega), \quad \phi = \gamma_0 \quad \text{on} \quad \Gamma_0, \quad \phi = \text{constant} \quad \text{on} \quad \Gamma_k, \quad \forall k = 1, \ldots, q \},
\] (6.20)
\[
W_g = \{ \phi | \phi \in H^1(\Omega), \quad \phi = \gamma_0 \quad \text{on} \quad \Gamma_0, \quad \phi = \gamma_k + \text{constant} \quad \text{on} \quad \Gamma_k, \quad \forall k = 1, \ldots, q \}.
\] (6.21)

It has to be understood that the constants occurring in (6.20), (6.21), are different according to $k$.

6.4.2. Description of the semi-implicit scheme

Let $\Delta t > 0$ be a time discretization step and let's denote by $\omega^n$ and $\psi^n$ approximate values of $\omega$ and $\psi$ at time $n \Delta t$. We approximate (6.14), (6.15) by
\[
\begin{align*}
\omega^0 &= \omega_0, \\
\lambda^0 &= \omega_0 |_\Gamma, \\
u^0 &= u_0;
\end{align*}
\] (6.22)
(6.23)
(6.24)
then for $n \geq 0$, assuming that $\omega^n$, $\lambda^n$, $u^n$ are known, we compute $\omega^{n+1}$, $\lambda^{n+1}$, $u^{n+1}$ as follows: solve
\[
\begin{align*}
\int_\Omega \frac{\omega^{n+1} - \omega^n}{\Delta t} \theta \, dx + &\nu \int_\Omega \nabla \omega^{n+1} \cdot \nabla \theta \, dx + \int_\Omega (u^n \cdot \nabla \omega^{n+1}) \theta \, dx \\
= &\int_\Omega \left( f_1^{n+1} \frac{\partial \theta}{\partial x_2} - f_2^{n+1} \frac{\partial \theta}{\partial x_1} \right) \, dx, \quad \forall \theta \in H^1_0(\Omega);
\end{align*}
\] (6.25)
\[
\omega^{n+1} = \lambda^n \quad \text{on} \quad \Gamma
\]
and then
\[
(\psi^{n+1} \in W_{g^{n+1}}, \forall \phi \in W_0, \text{ we have})
\]
\[
\int_\Omega \nabla \psi^{n+1} \cdot \nabla \phi \, dx = \int_\Omega \omega^{n+1} \phi \, dx - \int_\Gamma g^{n+1} \cdot \tau \phi \, d\Gamma.
\]  
(6.26)

To update \(\lambda^n\), solve
\[
\begin{cases}
\lambda^{n+1} \in M, \forall \mu \in M \text{ we have} \\
\int_\Omega \lambda^{n+1} \mu \, dx = \int_\Omega \nabla \psi^{n+1} \cdot \nabla \mu \, dx + \int_\Gamma g^{n+1} \cdot \tau \mu \, d\Gamma.
\end{cases}
\]
(6.27)

and finally, define \(u^{n+1}\) by
\[
u^{n+1} = \left\{ \frac{\partial \psi^{n+1}}{\partial x_2}, -\frac{\partial \psi^{n+1}}{\partial x_1} \right\}.
\]  
(6.28)

In (6.26), \(g^{n+1} = g((n+1)\Delta t)\) and \(W_{g^{n+1}}\) is the space defined by (6.21) when \(g = g^{n+1}\). In (6.27), \(M\) is a complementary subspace of \(H^1_0(\Omega)\) in \(H^1(\Omega)\) (i.e. \(H^1(\Omega) = H^1_0(\Omega) \oplus M\)).

Assuming that \(g^{n+1}, f^{n+1}\) and \(\lambda^n\) are sufficiently smooth, it is quite easy to show that problems (6.25) and (6.26) are well-posed (problem (6.27) is more delicate to analyze). Concerning problems (6.25), (6.26), we have the following:

**Proposition 6.1:** Suppose that \(\lambda^n, u^n, f^{n+1}\) and \(g^{n+1}\) are sufficiently smooth. Then problems (6.25) and (6.26) have a unique solution in \(H^1(\Omega)\) and \(W_{g^{n+1}}\), respectively.

**Proof:** Let’s discuss first the solution of the Dirichlet problem (6.25). From the properties of \(\lambda^n, u^n\) and \(f^{n+1}\), problem (6.25) will have a unique solution in \(H^1(\Omega)\) if we can prove that the bilinear functional
\[
\{ \theta, \theta' \} \rightarrow \int_\Omega \theta \theta' \, dx + \nu \Delta t \int_\Omega \nabla \theta \cdot \nabla \theta' \, dx + \Delta t \int_\Omega (u^n \cdot \nabla \theta) \theta' \, dx
\]
(6.29)
is \(H^1_0(\Omega)\)-elliptic, i.e. there exists an \(\alpha > 0\) such that
\[
\int_\Omega \theta^2 \, dx + \nu \Delta t \int_\Omega |\nabla \theta|^2 \, dx + \Delta t \int_\Omega (u^n \cdot \nabla \theta) \theta \, dx \geq \alpha \| \theta \|^2_{H^1_0(\Omega)}, \forall \theta \in H^1_0(\Omega).
\]  
(6.30)

Since \(\nabla \cdot u^n = 0\), we clearly have, \(\forall \theta, \theta' \in H^1_0(\Omega)\)
\[
\int_\Omega (u^n \cdot \nabla \theta) \theta' \, dx = - \int_\Omega (u^n \cdot \nabla \theta') \theta \, dx,
\]
which implies in turn that
\[
\int_\Omega (u^n \cdot \nabla \theta) \theta \, dx = 0, \forall \theta \in H^1_0(\Omega).
\]  
(6.31)
From (6.31), the bilinear functional defined by (6.29) is clearly \( H_0^1(\Omega) \)-elliptic (with, e.g., \( \alpha = \nu \Delta t \)). Since this functional is also continuous, the existence and uniqueness property follows from a simple variation of the Lax–Milgram Theorem (see, e.g., refs. [51] and [1, appendix 1] for more details on the solution of the nonhomogeneous Dirichlet problem by variational methods).

Let's consider now problem (6.26). Dropping the superscript \( n + 1 \) in (6.29), we introduce \( \psi_g \in W_g \) such that

\[
\psi_g = \gamma_k \quad \text{on} \quad \Gamma_k, \quad \forall k = 0, \ldots, q. \tag{6.32}
\]

If \( g \) is sufficiently smooth, \( W_g \) is not empty and therefore, such a \( \psi_g \) exists.

Now let \( \psi \) be the (hypothetical) solution of (6.26). Define \( \tilde{\psi} \) by

\[
\tilde{\psi} = \psi - \psi_g. \tag{6.33}
\]

We clearly have that \( \tilde{\psi} \) satisfies

\[
\begin{cases}
\tilde{\psi} \in W_0, \quad \forall \phi \in W_0, \quad \text{we have} \\
\int_{\Omega} \nabla \tilde{\psi} \cdot \nabla \phi \, dx = \int_{\Omega} \omega \phi \, dx - \int_{\Omega} \nabla \psi_g \cdot \nabla \phi \, dx - \int_{\Gamma} g \cdot \tau \phi \, d\Gamma.
\end{cases} \tag{6.34}
\]

The linear functional on the right-hand side of (6.34) is clearly continuous over \( H^1(\Omega) \) and therefore, over \( W_0 \). On the other hand, since \( \Omega \) is bounded and the functions of \( W_0 \) vanish on \( \Gamma_0 \), the bilinear functional

\[
\langle \phi, \phi' \rangle = \int_{\Omega} \nabla \phi \cdot \nabla \phi' \, dx \tag{6.35}
\]

is \( W_0 \)-elliptic, \( W_0 \) being equipped with the \( H^1(\Omega) \)-norm. Since the functional (6.35) is obviously continuous over \( W_0 \times W_0 \), it follows from the Lax–Milgram theorem that problem (6.34) has a unique solution, implying in turn that problem (6.26) also has a solution, possibly dependent on \( \psi_g \). In fact, the \( W_0 \)-ellipticity of the functional (6.35) implies the uniqueness of the solution of (6.26).

**Remark 2.1:** We can prove the existence of solutions of (6.25) (possibly not in \( H^1(\Omega) \)) with weaker assumptions on \( \lambda' \) than those in the statement of proposition 6.1, where we implicitly assumed that \( \lambda' \) was the trace of a function of \( H^1(\Omega) \).

**Remark 6.2:** The above techniques and results can be generalized to cases where \( \varphi \) and \( \partial \varphi / \partial n \) (respectively \( \partial \omega / \partial n \) and \( \partial \omega / \partial n \)) are given on a subset \( \Gamma_A \) (respectively \( \Gamma_B \)) of \( \Gamma \), such that \( \Gamma_A \cap \Gamma_B = \emptyset, \Gamma_A \cup \Gamma_B = \Gamma \).

### 6.5. A finite element method for the space discretization of problem (6.14), (6.15)

In this section, we shall assume that \( \Omega \) is a bounded, polygonal domain of \( \mathbb{R}^2 \). As in section 5, we denote by \( T_h \) a finite element triangulation of domain \( \Omega \). We associate with \( T_h \) the following
finite dimensional spaces:

\[ H^1_n = \{ \phi_h \mid \phi_h \in C^0(\bar{\Omega}), \quad \phi_h / T \in P_1, \quad \forall T \in T_h \}. \]

\[ H^1_{0h} = H^1_n \cap H^1_0(\Omega) = \{ \phi_h \mid \phi_h \in H^1_n, \quad \phi_h = 0 \text{ on } \Gamma \}. \]

In (6.36), \( P_1 \) is the space of the polynomials in \( x_1, x_2 \) of degree \( \leq l \) (in practice, we should use \( l = 1 \) or 2). We shall also need \( M_h \) defined by

\[ M_h = \{ \mu_h \mid \mu_h \in H^1_n, \quad \mu_h | T = 0, \quad \forall T \in T_h \} \quad \text{and} \quad \partial T \cap \Gamma = \emptyset. \]

We clearly have

\[ H^1_{0h} \oplus M_h = H^1_n. \]

We approximate then (6.14), (6.15) by the following set of coupled algebraic and differential equations (written in variational form):

\[ \int \frac{\partial \omega_h}{\partial t} \theta_h \, dx + \nu \int \nabla \omega_h \cdot \nabla \theta_h \, dx + \int \left( \frac{\partial \psi_h}{\partial x_2} \frac{\partial \omega_h}{\partial x_1} - \frac{\partial \psi_h}{\partial x_1} \frac{\partial \omega_h}{\partial x_2} \right) \theta_h \, dx \]

\[ = \int \left( f_{1h} \frac{\partial \theta_h}{\partial x_2} - f_{2h} \frac{\partial \theta_h}{\partial x_1} \right) \, dx, \quad \forall \theta_h \in H^1_{0h}, \]

\[ \omega_h(x, 0) = \omega_{0h}(x) \quad \text{(with } \omega_{0h} \in H^1_n), \]

\[ \int \nabla \psi_h \cdot \nabla \phi_h \, dx = \int \omega_h \phi_h \, dx - \int g_h \cdot \tau \phi_h \, d\Gamma, \]

\[ \partial \psi_h / \partial \tau = g_h \cdot n \quad \text{on } \Gamma, \]

\[ \psi_h(t), \omega_h(t) \in H^1_n. \]

In (6.40a)–(6.40d), \( f_{1h}, f_{2h}, \omega_{0h} \) and \( g_h \) are approximations of \( f_1, f_2, \omega_0 \) and \( g \), respectively, with \( g_h \) satisfying

\[ \int_{\Gamma_k} g_h \cdot n \, d\tau = 0, \quad \forall k = 0, 1, \ldots, q, \]

and compatible with the fact that \( \psi_h | T \subseteq P_1 \).

Using the approach of Section 6.4, we obtain a fully discrete approximation of problem (6.14), (6.15), defined as follows:

\[ \omega^0_h = \omega_{0h}, \]

\[ \lambda^0_h \in M_h, \quad \lambda^0_h = \omega_{0h} \quad \text{on } \Gamma, \]

\[ u^0_h = u_{0h}. \]
then for \( n \geq 0 \), assuming that \( \omega^n_h, \lambda^n_h, u^n_h \) are known, we compute \( \omega^{n+1}_h, \psi^{n+1}_h, \lambda^{n+1}_h, u^{n+1}_h \) as follows:

Solve:

\[
\begin{aligned}
\int_{\Omega} \frac{\omega^{n+1}_h - \omega^n_h}{\Delta t} \theta_h \, dx + \int_{\Omega} \nabla \omega^{n+1}_h \cdot \nabla \theta_h \, dx + \int_{\Omega} (u^n_h \cdot \nabla \omega^{n+1}_h) \theta_h \, dx \\
- \int_{\Omega} \left( f_{1h}^{n+1} \frac{\partial \theta_h}{\partial x_2} - f_{2h}^{n+1} \frac{\partial \theta_h}{\partial x_1} \right) \, dx, \quad \forall \theta_h \in H^1_{0h}, \\
\omega^{n+1}_h - \lambda^n_h &\in H^1_{0h},
\end{aligned}
\]  

(6.45)

and then

\[
\begin{aligned}
\psi^{n+1}_h &\in W^{\psi^{n-1}_h}; \quad \forall \phi_h \in W_{0h} \quad \text{we have} \\
\int_{\Omega} \nabla \psi^{n+1}_h \cdot \nabla \phi_h \, dx &= \int_{\Omega} \omega^{n+1}_h \phi_h \, dx - \int_{\Gamma} g^{n+1}_h \cdot \tau \phi_h \, d\Gamma.
\end{aligned}
\]  

(6.46)

To update \( \lambda^n_h \), solve

\[
\begin{aligned}
\lambda^{n+1}_h &\in M_h; \quad \forall \mu_h \in M_h \quad \text{we have} \\
\int_{\Omega} \lambda^{n+1}_h \mu_h \, dx &= \int_{\Omega} \nabla \psi^{n+1}_h \cdot \nabla \mu_h \, dx + \int_{\Gamma} g^{n+1}_h \cdot \tau \mu_h \, d\Gamma.
\end{aligned}
\]  

(6.47)

Finally, define \( u^{n+1}_h \) by

\[
u^{n+1}_h = \{ \partial \psi^{n+1}_h / \partial x_2, - \partial \psi^{n+1}_h / \partial x_1 \}. \quad \Box \]

(6.48)

In (6.44) (respectively (6.46)), \( u_{0h} \) is an approximation of the initial velocity \( u_0 \) (respectively \( g^{n+1}_h = g_h((n + 1)\Delta t) \)). Moreover, in (6.46), \( W_{0h} \) and \( W_{g^{n+1}_h} \) are defined by

\[
\begin{aligned}
W_{0h} &= \{ \phi_h | \phi_h \in H^1_h, \phi_h = 0 \quad \text{on} \quad \Gamma_k, \partial \phi_h / \partial \tau = 0 \quad \text{on} \quad \Gamma_k, \\
&\quad \forall k = 1, \ldots, q \},
\end{aligned}
\]  

(6.49)

\[
\begin{aligned}
W_{g^{n+1}_h} &= \{ \phi_h | \phi_h \in H^1_h, \phi_h = \gamma^{n+1}_0 \quad \text{on} \quad \Gamma_0, \partial \phi_h / \partial \tau = g^{n+1}_h \cdot n \\
&\quad \text{on} \quad \Gamma_k, \quad \forall k = 1, \ldots, q \},
\end{aligned}
\]  

(6.50)

with \( \gamma^{n+1}_0 \) defined by (6.16) (with \( g \) replaced by \( g^{n+1}_h \)).

Before discussing the solution of the linear problems (6.45), (6.46), (6.47), it is necessary to make some remarks concerning scheme (6.42)–(6.48).

Remark 6.3: In the particular case where \( l = 1 \) in (6.36) (piecewise linear approximations for \( \omega \).
and \( \psi \), it is quite convenient to replace the \( L^2 \)-scalar product occurring in (6.45), (6.46), (6.47) by the one (denoted \( (\cdot, \cdot)_h \)) defined by

\[
(\theta_h, \phi_h)_h = \frac{1}{2} \sum_{P \in \Sigma_h} m_P \theta_h(P) \phi_h(P), \quad \forall \theta_h, \phi_h \in H^1_h,
\]

(6.51)

where \( m_P \) is the measure of the polygonal union of those triangles of \( T_h \) with \( P \) as a common vertex. The scalar product (6.51) is clearly obtained from

\[
\int_{\Omega} \theta_h \phi_h \, dx = \sum_{T \in T_h} \int_T \theta_h \phi_h \, dx
\]

(6.52)

by applying the trapezoidal rule on each triangle on the right-hand side of (6.52). Similarly, we should use the trapezoidal rule to evaluate the boundary integrals in (6.46), (6.47).

For regular triangulations, applying the above numerical integration techniques goes back to classical finite difference schemes for discretizing the stream function–vorticity formulation of the Navier–Stokes equations.

Remark 6.4: Scheme (6.42)–(6.48) is of the semi-implicit type. Actually, it is quite easy to make it “more” implicit (and at the same time more stable and accurate), by using, instead of (6.45)–(6.48), the following iterative method:

\[
\lambda_h^{n+1,0} = \lambda_h^n, \quad u_h^{n+1,0} = u_h^n,
\]

(6.53)

then for \( m \geq 0 \), compute \( \omega_h^{n+1,m+1}, \psi_h^{n+1,m+1}, \lambda_h^{n+1,m+1}, u_h^{n+1,m+1} \) as follows:

Solve:

\[
\begin{align*}
\int_{\Omega} & \frac{\omega_h^{n+1,m+1} - \omega_h^n}{\Delta t} \theta_h \, dx + \nu \int_{\Omega} \nabla \omega_h^{n+1,m+1} \cdot \nabla \theta_h \, dx \\
& + \int_{\Omega} (u_h^{n+1,m} \cdot \nabla \omega_h^{n+1,m+1}) \theta_h \, dx = \int_{\Omega} \left( f_{1h}^{n+1} \frac{\partial \theta_h}{\partial x_2} - f_{2h}^{n+1} \frac{\partial \theta_h}{\partial x_1} \right) \, dx,
\end{align*}
\]

(6.54)

\( \forall \theta_h \in H^1_0 \),

\( \omega_h^{n+1,m+1} - \lambda_h^{n+1,m} \in H^1_0 \),

and then

\[
\begin{align*}
\psi_h^{n+1,m+1} & \in W_{k+1,h}; \quad \forall \phi_h \in W_0 \quad \text{we have} \\
\int_{\Omega} \nabla \psi_h^{n+1,m+1} \cdot \nabla \phi_h \, dx & = \int_{\Omega} \omega_h^{n+1,m+1} \phi_h \, dx - \int_{\Gamma} g_h^{n+1} \cdot \tau \mu_h \, d\Gamma.
\end{align*}
\]

(6.55)
To update $\lambda_{n+1,m}^{+1}$ solve

$$
\begin{align*}
\lambda_{n+1,m}^{+1} & \in M_h; \quad \forall \mu_h \in M_h \quad \text{we have} \\
\int_{\Omega} \nabla \psi_{n+1,m}^{+1} \cdot \nabla \mu_h \, dx + \int_{\Gamma} g_{n+1}^{+1} \cdot \tau \mu_h \, d\Gamma, 
\end{align*}
$$

(6.56)

and finally, define $u_{h}^{n+1,m+1}$ by

$$
u_{h}^{n+1,m+1} = \left\{ \frac{\partial \psi_{h}^{n+1,m+1}}{\partial x_2}, - \frac{\partial \psi_{h}^{n+1,m+1}}{\partial x_1} \right\}. \quad \Box
$$

(6.57)

Do $m = m + 1$ and go to (6.53).

The functions $\omega_{h}^{n+1}, \psi_{h}^{n+1}, \lambda_{h}^{n+1}, u_{h}^{n+1}$ will be obtained as the limits of $\omega_{h}^{n+1,m}, \psi_{h}^{n+1,m}, \lambda_{h}^{n+1,m}, u_{h}^{n+1,m}$ as $m \to +\infty$.

Some comments concerning the solution of the finite dimensional problems (6.45), (6.46), (6.47):

(Indeed, the following comments will also apply to the solution of problems (6.54), (6.55), (6.56)). Problems (6.45), (6.47) are classical finite dimensional variational problems. To solve them, we should substitute for $\theta_h$ (respectively $\mu_h$) the elements of a vector basis of $H_{0h}^1$ (respectively $M_h$) to construct equivalent linear systems. These systems can then be solved by classical direct or iterative methods. Problem (6.46) is less classical. To solve it, let’s define first $\Sigma_h$ (respectively $\Sigma_{0h}$) as the set of the vertices of $T_h$ (respectively, the vertices of $T_h$ not located on $\Gamma$). To each $P \in \Sigma_h$, associate then $w_P$ such that

$$
\begin{align*}
\begin{cases}
w_P \in H_h^1, \\
w_P(P) = 1; \quad w_P(Q) = 0, \quad \forall Q \in \Sigma_h, \quad Q \neq P.
\end{cases}
\end{align*}
$$

(6.58)

The following sets

$$
\begin{align*}
B_h = \{ w_P \}_P \in \Sigma_h, & \quad B_{0h} = \{ w_P \}_P \in \Sigma_{0h}, \\
B_h \setminus B_{0h} = \{ w_P \}_P \in \Sigma_h \setminus \Sigma_{0h}.
\end{align*}
$$

(6.59)

are vector bases of $H_h^1$, $H_{0h}^1$, $M_h$, respectively.

Finally, in order to construct a vector basis of $W_{0h}$, we associate to each $\Gamma_k$, $k = 1, \ldots, q$, a function $w_k'$ defined as follows:

$$
\begin{align*}
\begin{cases}
w_k' \in M_h, \quad w_k'|_{\Gamma_k} = 1, \\
w_k'|_{\Gamma_l} = 0, \quad \forall l \neq k, \quad 0 \leq l \leq q.
\end{cases}
\end{align*}
$$

(6.60)
It is quite obvious that

\[ \mathbf{B}_h^0 \cup \{ \mathbf{w}_k^q \}_{k=1}^q \]  

is a vector basis of \( \mathbf{W}_h^0 \). To compute \( \psi_h^{n+1} \) from (6.46), we first observe that (6.46) is equivalent to

\[
\begin{align*}
\int_{\Omega} \nabla \psi_h^{n+1} \cdot \nabla \mathbf{w}_P \, dx &= \int_{\Omega} \omega_h^{n+1} \mathbf{w}_P \, dx, \\
\forall \mathbf{w}_P \in \mathbf{W}_h^0,
\end{align*}
\]

(6.62a)

\[
\begin{align*}
\int_{\Omega} \nabla \psi_h^{n+1} \cdot \nabla \mathbf{w}_k' \, dx &= \int_{\Omega} \omega_h^{n+1} \mathbf{w}_k' \, dx - \int_{\Gamma_k} \mathbf{g}_h^{n+1} \cdot \mathbf{w}_k' \, d\Gamma, \\
\forall k = 1, \ldots, q,
\end{align*}
\]

(6.62b)

\[
\psi_h^{n+1} = \sum_{Q \in \mathbf{S}_h} \psi_h(Q) \mathbf{w}_Q + \sum_{l=1}^q C_l \mathbf{w}_l' + \sum_{l=0}^q \sum_{Q \in \mathbf{S}_h \cap \Gamma_l} \gamma_{lh}(Q) \mathbf{w}_Q,
\]

(6.62c)

where in the right-hand side of (6.62), we have dropped the superscript \( n + 1 \) and where \( \gamma_{lh} \) is defined by (6.16) (with \( \mathbf{g} \) replaced by \( \mathbf{g}_h^{n+1} \)).

By substituting the expansion (6.62c) of \( \psi_h^{n+1} \) into (6.62a) (6.62b) we obtain \( \{ \psi_h(Q) \}_{Q \in \mathbf{S}_h} \cup \{ C_l \}_{l=1}^q \) as the solution of a linear system whose matrix is symmetric and positive definite.

7. A backward method of characteristics for the incompressible Navier–Stokes equations. Implementation of downstream boundary conditions for flows in unbounded domains

7.1. Synopsis

Upwinding techniques to treat the advection term \( (\mathbf{u} \cdot \nabla)\mathbf{u} \) have always been quite popular among those solving the Navier–Stokes equations. For their finite difference (respectively, finite element) implementation see, e.g. refs. [11,12] (respectively, refs. [9,10,52]) and the references therein. The Petrov–Galerkin methods advocated by Hugues and his co-authors [53,54] also belong to this class of methods. It is, however, the opinion of the authors of this report that the most elegant way to “upwind” the Navier–Stokes equations is to use those backward methods of characteristics popularized by the investigators of the Laboratoire National d’Hydraulique (L.N.H.) at Electricite de France (see refs. [37,55] for further details and also ref. [56] for related techniques). In the following section 7.2, we shall briefly discuss a backward method of characteristics. Section 7.3 will be devoted to the finite element implementation of downstream boundary conditions which, from our numerical experiments, appear to be more absorbing than those of Dirichlet or Neumann types discussed in section 1.
7.2. Solution of the Navier–Stokes equations by a backward method of characteristics

Let’s define the (nonlinear) differential operator \( \frac{D\phi}{Dt} \) by

\[
\frac{D\phi}{Dt} = \frac{\partial \phi}{\partial t} + \sum_{j=1}^{N} u_j \frac{\partial \phi}{\partial x_j}, \quad \forall \phi \text{ sufficiently smooth},
\]

where \( \frac{D\phi}{Dt} \) is the Lagrangian time derivative of \( \phi \). With this notation, the momentum equation (1.1) takes the following form:

\[
\frac{Du}{Dt} - \nu \Delta u + \nabla p - f \quad \text{in } \Omega.
\]

Now, to \( \{x, t\} \in \Omega \times \mathbb{R}^+ \), we associate a flow \( X(x, t; \tau) \) (the characteristic flow) defined by

\[
\begin{cases}
\frac{dX}{d\tau} = u(X, \tau), \\
X(x, t; t) = x.
\end{cases}
\]

A simple time discretization of the Navier–Stokes equations using (7.2), (7.3) (i.e. obtained by integrating along the characteristic curves defined by (7.3)) is defined as follows:

\[
u^n = u_0,
\]

then for \( n \geq 0 \), assuming that \( u^n \) is known,

\[
\frac{u^{n+1}(x) - u^n(X^n(x; n \Delta t))}{\Delta t} - \nu \Delta u^{n+1}(x) + \nabla p^{n+1}(x) = f^{n+1}(x),
\]

\[
\nabla \cdot u^{n+1} = 0,
\]

\[
u^{n+1} = g^{n+1} \quad \text{on } \Gamma,
\]

where, in (7.4b), the flow \( X^n \) is defined as the solution of the (autonomous) differential system

\[
\begin{cases}
\frac{dX}{d\tau} = u^n(X), \\
X(x; (n + 1) \Delta t) = x.
\end{cases}
\]

The Stokes problem (7.4b)–(7.4d) can be solved by the methods described in section 4.

Remark 7.1: Eq. (7.4b) is slightly incorrect. Actually, if the curve described by \( X^n(x; \tau) \) for \( \tau \in [n\Delta t, (n + 1)\Delta t] \), is contained in \( \Omega \), we shall use (7.4b) to discretize (1.1) at time \( (n + 1)\Delta t \). But, if there exists \( \tau^*_n, n\Delta t < \tau^*_n < (n + 1)\Delta t \) such that

\[
X^n(x; \tau^*_n) \in \Gamma,
\]
we shall discretize (1.1) at \( \{x, (n + 1)\Delta t\} \) by

\[
\frac{\mathbf{u}_n^{n+1}(x) - \mathbf{u}_n^n(X^n(x; \tau^*_n), \tau^*_n)}{(n + 1) \Delta t} - \nu \frac{\Delta \mathbf{u}_n^{n+1}(x)}{\tau^*_n} + \nabla p_n^{n+1}(x) = f_n^{n+1}(x). \tag{7.6}
\]

The finite element implementation of the above method is not difficult, in principle, particularly with the finite element approximation described in section 5 (see refs. [37,55,56] for more details and the results of numerical experiments).

### 7.3. Finite element implementation of downstream boundary conditions for flows in unbounded domains

One of the major problems associated with the numerical simulation of incompressible viscous flows (at high Reynolds numbers particularly) is to find a convenient set of downstream boundary conditions when the physical unbounded domain is replaced by a bounded computational one.

A partial answer to this question can be found in ref. [57], but the methods discussed in this last reference still have to be further analyzed in view of their computational implementation. In this section, we shall discuss some downstream boundary conditions whose practical implementation is quite easy through the finite element approximation introduced in section 5.

For flows at high Reynolds numbers (which is one of our main concerns), several authors have suggested to take as downstream boundary conditions the following inviscid momentum equation (where for simplicity we have taken \( f = 0 \))

\[
\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} + \nabla p = 0. \tag{7.7}
\]

Let's describe the practical implementation of condition (7.7) in the particular case of the flow problem associated with fig. 1.1 of section 1.1. Suppose that a finite element approximation \( \{\mathbf{u}_h^n, p_h^n\} \) of \( \{\mathbf{u}, p\} \) is known over \( \overline{\Omega} \) at time \( n\Delta t \). We should predict \( \mathbf{u}_h^{n+1} \) over \( \Gamma_\infty \) as follows:

1. We first compute the trace \( \mathbf{u}_h^{n+1} \) of \( \mathbf{u}_h^n \) defined by:

   \[
   \gamma \mathbf{v}_h \cdot \mathbf{v}_h \Delta t \mathbf{v}_h \cdot \mathbf{v}_h = \int_{\Omega} \mathbf{u}_h^n \cdot \nabla \mathbf{u}_h^n \cdot \mathbf{v}_h \, dx + \int_{\Omega} \nabla p_h^n \cdot \mathbf{v}_h \, dx = 0, \tag{7.8}
   \]

   \[\forall \mathbf{v}_h \in V_h^+; \quad \mathbf{u}_h^{n+1} \in V_h, \quad \mathbf{u}_h^{n+1} = u_\infty((n + 1) \Delta t) \quad \text{on} \quad \Gamma_\infty,\]

   where

   \[V_h^+ = \{ \mathbf{v}_h | \mathbf{v}_h \in V_h, \mathbf{v}_h = 0 \quad \text{on} \quad \Gamma_\infty \}.\]

   In practice, we shall use numerical integration (trapezoidal rule for example) to approximate the first integral in (7.8), namely

   \[
   \int_{\Omega} \frac{\mathbf{u}_h^{n+1} - \mathbf{u}_h^n}{\Delta t} \cdot \mathbf{v}_h \, dx,
   \]

   * Derived from section 5.
in such a way that the corresponding mass matrix will be diagonal. Then we shall take for \( v_h \) those basis functions of \( V_h \) associated with the nodes of \( T_h \) (vertices and mid-points if \( V_h \) is defined by either (5.2) or (5.4); vertices only if \( V_h \) is defined by (5.5)) belonging to \( \Gamma^+_\infty \) (interior of \( \Gamma^+_\infty \)).

(ii) Once the trace \( \tilde{g}^{n+1} \) of \( \tilde{u}^{n+1} \) is known, we obtain \( g_h^{n+1} \) by correcting \( \tilde{g}^{n+1} \) in such a way that

\[
\int_{\Gamma} \tilde{g}^{n+1} \cdot n \, d\Gamma = 0. \tag{7.9}
\]

Such a correction can be obtained by the following variant of (5.12) (see section 5.2):

\[
g_h^{n+1} = \tilde{g}^{n+1} - \left( \frac{\int_{\Gamma} \tilde{g}^{n+1} \cdot n \, d\Gamma}{\int_{\Gamma} n \cdot n_h \, d\Gamma} \right) n_h, \tag{7.10}
\]

with \( n_h \) still defined by (5.11). In order to obtain relation (7.9), we can also use variants of (7.10) in which the correction of \( \tilde{g}^{n+1} \) is only done on \( \Gamma^+_\infty \).

The above method to predict \( u^{n+1} \) on \( \Gamma^+_\infty \) can easily be modified to be combined with the operator splitting methods discussed in sections 2.3 and 5.4.

The above technique has been tested for flows around cylinders and the corresponding numerical results will be reported in ref. [38], together with comparisons to other methods. From these first experiments it seems that these new boundary conditions are more “transparent” than the one obtained by just prescribing \( u = u_\infty \) on \( \Gamma^+_\infty \), or by mixing, as suggested by some authors, Dirichlet et Neumann boundary conditions. Actually, it seems that re-introducing some viscosity in (7.7) may improve the quality of the numerical results. Such modification of the technique described just above will be discussed in ref. [38], where associated numerical results will also be shown. (We can also take advantage of the method of characteristics to implement downstream boundary conditions with good absorbing properties; see, again, ref. [38] for details.)

8. Numerical simulation of some incompressible viscous flows

The main goal of this section is to present the results of numerical experiments concerning the simulation of two and three dimensional incompressible viscous flows. In section 8.1 we shall present the results of a plane jet simulation, and then in sections 8.2 and 8.3 the simulation of flows around and inside air intakes at high angle of attack.

8.1. Plane jet simulation

We consider a plane jet problem in which the flow domain \( \Omega \) is the right half space

\[
(x \mid x = (x_i)_{i=1}^2, \quad x_1 > 0, \quad x_2 \in \mathbb{R}).
\]

We assume that the fluid viscosity \( \nu \) is equal to 1/2000, that the fluid is at rest at time \( t = 0 \), i.e.

\[
u(x, 0) = 0, \quad \forall x \in \Omega,
\]
that the jet aperture is located on the line $x_1 = 0$, is centered at $x = (0, 0)$ and is $1/16$ wide. We suppose also that the jet profile is parabolic, with a maximal velocity equal to 1, the jet being horizontal at $x_1 = 0$ (we have then $Re = 125$).

We have used several computational domains, all defined by $]-L/2, L/2[ \times ]0, L[$, with $L$ taking the values 1, 2 and $e^2$, successively ($e = 2.718\ldots$).

The flow simulation was done using stream function vorticity formulation of section 6, combined to a fully implicit (second order in time) variant of the scheme discussed in sections 6.4 and 6.5; the finite element discretization was achieved using piecewise linear approximations (i.e. $l = 1$ in (6.36)) for $\omega$ and $\psi$. The details of such an approximation may be found in ref. [50].
Taking $\Delta t = 0.1$, the following results have been obtained:

(i) Figs. 8.1 to 8.8 ($t = 0.1, 0.5, 0.7, 0.9, 1.2, 1.5, 2.5, 4$) show the flow pattern (particularly the vortex creation) close to the jet aperture at different times. For these results we have been using $L = 1$ and a uniform finite element triangulation of 2048 triangles.

(ii) Since at $t \approx 4$ the computational domain appears to be too small, we have been using $L = 2$ and a uniform finite element triangulation of 8192 triangles.

The corresponding results are reported in figs. 8.9 to 8.12 ($t = 4.5, 8.5, 8.9, 9.5$). These figures show some interesting phenomena such as a symmetry breaking occurring close to $t = 8.5$ and the creation of small eddies in the lower shear layer between $t = 8.5$ and 8.9. These secondary
eddy remains if we decrease $\Delta t$ and the space discretization step, and if we increase the size of
the computational domain.

(iii) In order to check the influence of the mesh and of the size of the computational domain we
have been running a numerical simulation with $L = e^2$ and the nonuniform finite element grid of
fig. 8.13 (18432 triangles). The results confirm those of (i) and (ii): in Fig. 8.14, we have shown
the flow pattern corresponding to $t = 13$.

8.2. Flow in a nozzle at high incidence

The second test problem that we consider is much more complicated than the first one. It
concerns the simulation of an incompressible viscous flow inside and around a (two-dimensional)
EXPONENTIAL GRID $N=M=95$

Fig. 8.13.

STREAM FCN. $\psi$  $T=13.0$

Fig. 8.14.
nozzle at high incidence (40 degrees) and at Re = 750 (the characteristic length is taken as the
distance between the walls of the nozzle). We used the same type of finite element approxima-
tions as in section 5, with V_h defined by (5.4). Figs. 8.15 and 8.16 show the details of the
triangulations T_h and T_h, respectively, close to the air intake. Using Δt = 0.05, we obtained the
streamlines and vortex pattern shown in figs. 8.17 to 8.21 (t = 0., 2., 4., 6., 8.). The initial velocity
distribution is associated to the corresponding steady Stokes flow and a suction phenomenon is
simulated inside the nozzle. We have shown in fig. 8.22 the visualization of experiments done at
ONERA by Werle [89] for a Reynolds number of same order of magnitude and for the same
angle of attack. We observe the good agreement between the computed and experimental results.
In figs. 8.23 to 8.25 we have shown a color visualisation (with the flow coming from the left hand
side) of the vorticity at t = 4., 6., 8.

A similar simulation has been done for Re = 1500. A finer grid was required for these
calculations as shown in figs. 8.26 (pressure grid) and 8.27 (velocity grid). Using Δt = 0.025 we
obtained the velocity configuration of figs. 8.28 (t = 3.5) and 8.29 (t = 4.5); a visualisation of the
vorticity at the same time cycles is shown in figs. 8.30 and 8.31.

8.3. Flow in a three-dimensional nozzle at high incidence

We consider here a flow inside and around a three-dimensional nozzle at Re = 200, for an
angle of attack of 40 degrees. The corresponding simulations have been done using the θ-scheme
(2.34)–(2.37) of section 2.3.2, with $\theta = 1 - \sqrt{2}/2$, combined to the three-dimensional variant of the finite element approximation of section 5 defined by (5.5) (bubble-function for the velocity, piecewise linear pressure). Indeed, a first calculation has been done during 30 time cycles (with $\Delta t = 0.2$), using the grid shown in fig. 8.32 (3096 nodes, 16632 tetrahedra). From these results the grid is locally refined (see fig. 8.33), using the techniques of section 15 (we have now 4632 nodes and 24457 tetrahedra), and the integration is performed with this new grid during 20 time cycles with the same $\Delta t$. The results obtained at $t = 10$, without and with refinement, are shown in figs. 8.34 and 8.35. (Figs. 8.32, 8.33 on page 138, figs. 8.34, 8.35 on page 137.)

We observe an improved accuracy for the pressure and vorticity with the refined grid.
Fig. 8.17. $Re = 750$: $t = 0$. 
Fig. 8.18. Re = 750; $t = 2$. 
Fig. 8.19. $Re = 750$; $t = 4$. 
Fig. 8.20. \( \text{Re} = 750; \ t = 6. \)
Fig. 8.21. \( Re = 750; \ t = 8. \)
Fig. 8.22. ONERA experimental results: courtesy of DRET and H. Werle.
Fig. 8.23. Vorticity visualization, Re = 750; $t = 4$.

Fig. 8.24. Vorticity visualization, Re = 750; $t = 6$.

Fig. 8.25. Vorticity visualization, Re = 750; $t = 8$. 
Fig. 8.26. Pressure grid.
Fig. 8.27. Velocity grid.
Fig. 8.28. Velocity visualization, Re = 1500; t = 3.5.
Fig. 8.29. Velocity visualization, Re = 1500; t = 4.5.

Fig. 8.30. Vorticity contours, Re = 1500; t = 3.5.

Fig. 8.31. Vorticity contours, Re = 1500; t = 4.5.
Fig. 8.34. Vorticity distribution for the non-refined grid.

Fig. 8.35. Vorticity distribution for the refined grid.
Fig. 8.32. Non-refined grid.

Fig. 8.33. Refined grid by self-adaptive techniques.
PART II: NUMERICAL SIMULATION OF COMPRESSIBLE VISCous FLOWS

9. Generalities. Synopsis

The numerical simulation of compressible viscous flows has been the subject of quite intensive efforts these last years. One of the main motivations of this fact is the computer (and more recently supercomputer) design of advanced aircrafts and space vehicles.

The main goal of this part II is to discuss finite element methods and associated algorithms, that we have used these last years to solve various compressible viscous flow problems, many of them being of industrial interest.

The numerical techniques to be discussed below are indeed natural extensions of those discussed in part I for incompressible viscous flows. However we have included in this part II a brief discussion of a self-adaptive grid refinement method that we found most useful in order to increase the local accuracy in those regions where the solution exhibits very strong gradients.

A large part of the material discussed in the following sections is contained in refs. [58,59].

10. Formulation of the Navier–Stokes equations for compressible viscous fluids

Let \( \Omega \subset \mathbb{R}^N \) (\( N = 2, 3 \) in practice) be the flow domain and \( \Gamma \) be its boundary. The non-dimensional conservative form of the equations is given below by:

\[
\begin{align*}
\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{u}) &= 0, \\
\frac{\partial \rho \mathbf{u}}{\partial t} + \nabla \cdot (\rho \mathbf{u} \otimes \mathbf{u}) + \nabla p &= \frac{1}{\text{Re}} \left[ \Delta \mathbf{u} + \frac{1}{\text{Re}} \nabla (\nabla \cdot \mathbf{u}) \right], \\
\frac{\partial \varepsilon}{\partial t} + \nabla \cdot (\varepsilon \mathbf{u} + p \mathbf{u}) &= \frac{1}{\text{Re}} \left[ \nabla \cdot \left( \mathbf{u} \left( -\frac{2}{3} \nabla \cdot \mathbf{u} + \nabla \mathbf{u} + \nabla \mathbf{u}^\top \right) \right) + \frac{\gamma}{\text{Pr}} \frac{\varepsilon}{\Delta \varepsilon} \right];
\end{align*}
\]

in (10.1), we have normalized by the subscript \( \text{r} \)

(i) the density \( \rho \) by \( \rho_r \),
(ii) the velocity \( \mathbf{u} \) by \( |\mathbf{u}_r| \),
(iii) the specific internal energy \( \varepsilon \) by \( |\mathbf{u}_r|^2 \),
(iv) the pressure \( p \) by \( \rho_r |\mathbf{u}_r|^2 \).

The pressure obeys the ideal gas law:

\[ p = (\gamma - 1) \rho \varepsilon \]  

Finally, we have also normalized the total energy \( \varepsilon \) by \( \rho_r |\mathbf{u}_r|^2 \), and we have

\[ \varepsilon = \rho \varepsilon + p |\mathbf{u}|^2 / 2. \]

The constants \( \text{Re}, \text{Pr} \) and \( \gamma \) are the Reynolds number, the Prandtl number and the ratio of specific heats, respectively (\( \gamma = 1.4 \) in air).
The above equations express the conservation of mass, momentum and energy. We normalize the temperature $T$ by $|\mathbf{u}_r|^2/c_\gamma$, implying that

$$T = \epsilon. \quad (10.4)$$

From (10.1)–(10.4), we can deduce the following non-conservative form of the Navier-Stokes equations:

$$\frac{\partial \rho}{\partial t} + \mathbf{u} \cdot \nabla \rho + \rho \nabla \cdot \mathbf{u} = 0, \quad (10.5)$$

$$\rho \frac{\partial \mathbf{u}}{\partial t} + \rho (\mathbf{u} \cdot \nabla) \mathbf{u} + (\gamma - 1) \nabla (\rho T) = \frac{1}{\text{Re}} \left[ \Delta \mathbf{u} + \frac{1}{2} \nabla (\nabla \cdot \mathbf{u}) \right], \quad (10.6)$$

$$\rho \frac{\partial T}{\partial t} + \rho \mathbf{u} \cdot \nabla T + (\gamma - 1) \rho T \nabla \cdot \mathbf{u} = \frac{1}{\text{Pr}} \left[ \frac{\gamma}{\text{Pr}} \Delta T + F(\nabla \mathbf{u}) \right], \quad (10.7)$$

where (10.2)–(10.4) still hold.

For three-dimensional flows, we have $\mathbf{u} = \{ u, v, w \}$ and $F(\cdot)$ in (10.7) has the following expression:

$$F(\nabla \mathbf{u}) = \frac{4}{3} \left[ \left( \frac{\partial u}{\partial x} \right)^2 + \left( \frac{\partial v}{\partial y} \right)^2 + \left( \frac{\partial w}{\partial z} \right)^2 \right] + \left( \frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} \right)^2 + \left( \frac{\partial w}{\partial x} + \frac{\partial u}{\partial z} \right)^2 + \left( \frac{\partial v}{\partial z} + \frac{\partial w}{\partial y} \right)^2$$

$$- \frac{4}{3} \left( \frac{\partial u}{\partial x} \frac{\partial v}{\partial y} + \frac{\partial u}{\partial x} \frac{\partial w}{\partial z} + \frac{\partial v}{\partial y} \frac{\partial w}{\partial z} \right). \quad (10.8)$$

We consider external flows around airfoils or inlets for instance: the domain of computation is described in fig. 10.1. Let $\Gamma_\infty$ be the far-field boundary of the domain; we introduce then

$$\Gamma_\infty^- = \{ x | x \in \Gamma_\infty, \mathbf{u}_\infty \cdot \mathbf{n} < 0 \}, \quad (10.9)$$

$$\Gamma_\infty^+ = \Gamma_\infty \setminus \Gamma_\infty^-, \quad (10.10)$$

where $\mathbf{u}_\infty$ denotes the free stream velocity and $\mathbf{n}$ the unit vector of the outward normal to $\Gamma$. We assume the flow to be uniform at infinity, and the corresponding variables to be normalized by the free stream values; then for example, we prescribe at infinity

$$\mathbf{u} = \mathbf{u}_\infty = \begin{pmatrix} \cos \alpha \\ \sin \alpha \end{pmatrix}, \quad \alpha \text{ angle of attack}, \quad (10.11)$$

$$\rho = 1, \quad (10.12)$$

$$T = T_\infty = 1/\gamma (\gamma - 1) M_\infty^2, \quad (10.13)$$

where $M_\infty$ denotes the free stream Mach number.
The problem of the boundary conditions on the computational boundary $\Gamma_\infty$ is more delicate and will be discussed later.

On the rigid boundary $\Gamma_B$, we shall use the following conditions:

$$\mathbf{u} = \mathbf{0} \quad \text{(no-slip condition)},$$  \hspace{1cm} (10.14)

$$T = T_B = T_\infty (1 + (\gamma - 1)/2) M_\infty^2 \quad \text{(free stream total temperature)}. \hspace{1cm} (10.15)$$

The value of the pressure (and density) at a rigid boundary is deduced from the momentum equation. This operation is done implicitly through a generalized Stokes solver (to be discussed hereafter). An alternative is to combine (10.14) and (10.1); we obtain

$$\frac{\partial p}{\partial n} = \frac{1}{\text{Re}} \left[ \Delta \mathbf{u} + \frac{1}{3} \nabla (\nabla \cdot \mathbf{u}) \right] \cdot \mathbf{n}. \hspace{1cm} (10.16)$$

which can be used through an explicit procedure.

Finally, since we consider time dependent equations (even if we are looking for steady solutions), initial conditions have to be added; we shall take

$$\rho(x, 0) = \rho_0(x), \hspace{1cm} (10.17)$$

$$\mathbf{u}(x, 0) = \mathbf{u}_0(x), \hspace{1cm} (10.18)$$

$$T(x, 0) = T_0(x). \hspace{1cm} (10.19)$$

Solving the compressible Navier–Stokes equations is a most difficult task. Most of the existing numerical solution methods are based on Finite Differences techniques, for both space and time discretizations (see refs. [11,60–63] and the references therein), Spectral Methods have also been considered (see refs. [64,65]).

More recently, Finite Element Methods have been investigated to solve the above problem (see refs. [66,67]). As a final comment, we would like to mention ref. [68] where some theoretical results concerning the compressible Navier–Stokes equations have been proved, assuming that the data are “small”. The above reference list is far from being complete.
11. Application of operator splitting methods to the solution of the time-dependent compressible Navier–Stokes equations

11.1. Transformation of the equations

Using convenient operator splitting methods described in part I, section 2, for the time discretization of the compressible Navier–Stokes equations, we are able to decouple the difficulties due to the nonlinearity and to the pressure computation at rigid boundaries, respectively.

In order to extend the methodology previously used for solving the incompressible Navier–Stokes equations (see part I), we introduce the following new variable:

\[ \sigma = \ln \rho. \]  

(11.1)

With this new variable, the system (10.5)–(10.7) becomes:

\[ \frac{\partial \sigma}{\partial t} + \nabla \cdot u + u \cdot \nabla \sigma = 0, \]  

(11.2)

\[ \frac{\partial u}{\partial t} + (u \cdot \nabla)u + (\gamma - 1)(T \nabla \sigma + \nabla T) = \frac{e^{-\sigma}}{Re} (\Delta u + \frac{1}{2} \nabla (\nabla \cdot u)). \]  

(11.3)

\[ \frac{\partial T}{\partial t} + u \cdot \nabla T + (\gamma - 1)T \nabla \cdot u = \frac{e^{-\sigma}}{Re} \left( \frac{\gamma}{Pr} \Delta T + F(\nabla u) \right); \]  

(11.4)

the eqs. (11.3), (11.4) can also be written as:

\[ \frac{\partial u}{\partial t} - \mu \Delta u + \beta \nabla \sigma - \Psi(\sigma, u, T) = 0, \]  

(11.5)

\[ \frac{\partial T}{\partial t} - \Pi \Delta T - \chi(\sigma, u, T) = 0, \]  

(11.6)

with \( \delta \): a mean value of the inverse of the density, \( \delta = 1 \) is a possible value.

\[ \nu = 1/Re, \quad \mu = \nu \delta, \quad \Pi = \nu \delta /Pr, \]

\[ \beta = (\gamma - 1)T_B = \frac{1}{\gamma} \left( \frac{\gamma - 1}{2} + \frac{1}{M_{\infty}^2} \right), \]

\[ \Psi(\sigma, u, T) = - (\gamma - 1) \left[ \nabla T + (T - T_B) \nabla \sigma \right] - (u \cdot \nabla)u + \nu \left[ e^{-\sigma}(\Delta u + \frac{1}{2} \nabla (\nabla \cdot u)) - \Delta u \right]. \]

\[ \chi(\sigma, u, T) = - (\gamma - 1) T \nabla \cdot u - u \cdot \nabla T + \frac{\gamma \nu}{Pr} (e^{-\sigma} \Delta T - \delta \Delta T) + \nu e^{-\sigma} F(\nabla u). \]
11.2. An operator splitting method

Let $\Delta t(>0)$ be a time discretization step and $\theta$ belongs to $]0, 1/2[$. This method is derived from scheme (2.6)-(2.9) and is obtained as follows:

\[
\begin{aligned}
\sigma^0 &= \sigma_0 = \ln \rho_0, \\
\mathbf{u}^0 &= \mathbf{u}_0, \\
T^0 &= T_0,
\end{aligned}
\]  
(11.7)

then for $n \geq 0$, starting from $\{\sigma^n, \mathbf{u}^n, T^n\}$, we solve

\[
\begin{aligned}
\frac{\sigma^{n+1} - \sigma^n}{\theta \Delta t} + \nabla \cdot \mathbf{u}^{n+1} &= -\mathbf{u}^n \cdot \nabla \sigma^n, \\
\frac{\mathbf{u}^{n+1} - \mathbf{u}^n}{\theta \Delta t} - a \mu \Delta \mathbf{u}^{n+\theta} + b \nu \sigma^{n+\theta} &= \Psi(\sigma^n, \mathbf{u}^n, T^n) + b \mu \Delta \mathbf{u}^n, \\
\frac{T^{n+1} - T^n}{\theta \Delta t} - a \Pi \Delta T^{n+\theta} &= \chi(\sigma^n, \mathbf{u}^n, T^n) + b \Pi \Delta T^n,
\end{aligned}
\]  
(11.8)

\[
\begin{aligned}
\frac{\sigma^{n+1} - \sigma^n}{(1 - 2\theta) \Delta t} + \mathbf{u}^{n+1} \cdot \nabla \sigma^{n+1} &= -\nabla \cdot \mathbf{u}^{n+1}, \\
\frac{\mathbf{u}^{n+1} - \mathbf{u}^n}{(1 - 2\theta) \Delta t} - b \mu \Delta \mathbf{u}^{n+1} - \Psi(\sigma^{n+1}, \mathbf{u}^{n+1}, T^{n+1}) &= a \mu \Delta \mathbf{u}^{n+\theta} - b \nu \sigma^{n+\theta}, \\
\frac{T^{n+1} - T^n}{(1 - 2\theta) \Delta t} - b \Pi \Delta T^{n+1} - \chi(\sigma^{n+1}, \mathbf{u}^{n+1}, T^{n+1}) &= a \Pi \Delta T^{n+\theta},
\end{aligned}
\]  
(11.9)

\[
\begin{aligned}
\frac{\sigma^{n+1} - \sigma^{n+1}}{\theta \Delta t} + \nabla \cdot \mathbf{u}^{n+1} &= -\mathbf{u}^{n+1} \cdot \nabla \sigma^{n+1}, \\
\frac{\mathbf{u}^{n+1} - \mathbf{u}^n}{\theta \Delta t} - a \mu \Delta \mathbf{u}^{n+1} + b \nu \sigma^{n+1} &= \Psi(\sigma^{n+1}, \mathbf{u}^{n+1}, T^{n+1}) + b \mu \Delta \mathbf{u}^{n+1}, \\
\frac{T^{n+1} - T^n}{\theta \Delta t} - a \mu \Delta T^{n+1} &= \chi(\sigma^{n+1}, \mathbf{u}^{n+1}, T^{n+1}) + b \Pi \Delta T^{n+1},
\end{aligned}
\]  
(11.10)

with $0 < a, b < 1, a + b = 1$, satisfying (2.28) ($a, b$ play the role of $\alpha, \beta$ in the previous part). The boundary conditions will be specified in the following.
11.3. Some comments and remarks concerning scheme (11.8)–(11.10)

Using the above operator splitting method, we have been able to decouple nonlinearity and pressure (and density) computation at rigid boundaries. We shall describe, in the following sections, the specific treatment of the subproblems encountered at each step.

The linear systems (11.8), (11.10) are generalized Stokes problems which are also found in the shallow-water equations [69]. We use a direct Stokes solver (section 12) introduced by Glowinski–Pironneau [70], which reduces the problem to a cascade of Poisson problems.

For the nonlinear system (11.9), we introduce either a rational Runge–Kutta technique using dynamic relaxation [71] or a least squares formulation and a preconditioned conjugate gradient algorithm (section 13).

12. Solution of the generalized Stokes subproblems

12.1. Generalities. Synopsis

At each full step of the splitting method (11.8)–(11.10), we have to solve two linear problems of the following type

\[ \alpha \sigma + \nabla \cdot \mathbf{u} = g, \quad (12.1) \]

\[ \alpha \mathbf{u} - a \mu \Delta \mathbf{u} + \beta \nabla \sigma = f, \quad (12.2) \]

\[ \alpha T - a \Pi \Delta T = h, \quad (12.3) \]

\[
\begin{cases}
\sigma = 0 & \text{on } \Gamma_-^-, \\
\partial \sigma / \partial n = 0 & \text{on } \Gamma_+^-, \\
\mathbf{u} = \mathbf{0} & \text{on } \Gamma_B, \\
\mathbf{u} = \mathbf{u}_\infty & \text{on } \Gamma_-^+, \\
\partial \mathbf{u} / \partial n = \mathbf{0} & \text{on } \Gamma_+^+, \\
T = T_B & \text{on } \Gamma_B, \\
T = T_\infty & \text{on } \Gamma_-^+, \\
\partial T / \partial n = 0 & \text{on } \Gamma_+^+. 
\end{cases} \quad (12.4)\]

where \( \alpha \) is a positive parameter and where \( g, f, h \) are given functions.

The temperature \( T \) is the solution of the decoupled Poisson equation (12.3), (12.6). We shall describe below a direct method for solving the generalized Stokes problem satisfied by \( \{ \sigma, \mathbf{u} \} \).

A direct method for solving Stokes problem has been introduced by Glowinski and Pironneau [70] and has been extensively used as a component of an operator splitting method for the
solution of the time dependent incompressible Navier–Stokes equations \[1, 38\] in place of the iterative methods described in section 4. This Stokes solver has been generalized by Benque et al. \[69, 72\] in order to solve the shallow water equations; problem (12.1), (12.2), (12.4), (12.5) is of the same type and therefore can be solved by techniques very close to those used in refs. \[69, 72\].

The proposed method reduces the solution of the problem (12.1), (12.2), (12.4), (12.5) to the solution of a finite number of Poisson problems plus the solution of a boundary integral equation.

Let us introduce the motivation of the algorithm; for more precisions and theoretical results, see refs. \[1, 70\]. First we apply the operator \( \nabla \cdot \) to eq. (12.2), we obtain:

\[
\alpha \nabla \cdot u - \alpha \mu \Delta (\nabla \cdot u) + \beta \Delta \sigma = \nabla \cdot f; \tag{12.7}
\]

on the other hand, eq. (12.1) yields

\[
\nabla \cdot u = g - \alpha \sigma. \tag{12.8}
\]

Then, eq. (12.7) can be written with only \( \sigma \) as unknown variable:

\[
\alpha^2 \sigma - (\beta + \alpha \mu) \Delta \sigma = \alpha g - \nabla \cdot f - \alpha \mu \Delta g, \tag{12.9}
\]

with the boundary conditions (12.4).

In order to have a well-posed problem in \( \sigma \), it is necessary to have an additional boundary condition of type:

\[
\sigma = \lambda \quad \text{on} \quad \Gamma_B.
\]

If \( \sigma \) is evaluated, then the velocity \( u \) is solution of

\[
\alpha u - \alpha \mu \Delta u = -\beta \nabla \sigma + f, \tag{12.10}
\]

with the boundary conditions (12.5).

The value of \( \lambda \) has to be calculated in order that eq. (12.1) is satisfied. Therefore we introduce the function \( \psi \) solution of:

\[
\alpha^2 \psi - \zeta \Delta \psi = \alpha \sigma + \nabla \cdot u - g, \tag{12.11}
\]

\[
\begin{align*}
\psi &= 0 \quad \text{on} \quad \Gamma_B \cup \Gamma^-_\infty, \\
\frac{\partial \psi}{\partial n} &= 0 \quad \text{on} \quad \Gamma^+_\infty
\end{align*} \tag{12.12}
\]

with \( \zeta = \beta + \alpha \mu \).

It can easily be proved using (12.1), (12.2) that the function \( \psi \) satisfies an equation of the type:

\[
\mathcal{A} \psi - \mathcal{B} \Delta \psi + \mathcal{C} \Delta^2 \psi = 0, \quad \mathcal{A}, \mathcal{B}, \mathcal{C} > 0.
\]

If we can prove that \( \psi \equiv 0 \), then eq. (12.1) will be satisfied. The function \( \psi \) depends linearly of \( \lambda \), and the following boundary conditions have to be satisfied

\[
\begin{align*}
\frac{\partial \psi}{\partial n} &= 0 \quad \text{on} \quad \Gamma_B \cup \Gamma^-_\infty, \\
\psi &= 0 \quad \text{on} \quad \Gamma^+_\infty
\end{align*} \tag{12.13}
\]
therefore we introduce the symmetric linear operator

\[ A : \lambda \rightarrow \left\{ \partial \psi / \partial n \mid \Gamma_B \cup \Gamma_\infty, \quad \psi \mid \Gamma_\infty \right\}. \]

Then the solution method can be written as follows:

**Construction of the implicit integral operator A**

Let us define \( \sigma_\lambda, u_\lambda, \psi_\lambda \) solutions of (12.14), (12.15), (12.16) respectively

\[
\begin{cases}
\alpha^2 \sigma_\lambda - \xi \Delta \sigma_\lambda = 0, \\
\sigma_\lambda = \lambda \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty, \\
\partial \sigma_\lambda / \partial n = \lambda \quad \text{on} \quad \Gamma_\infty^+, \\
\end{cases} \tag{12.14}
\]

\[
\begin{cases}
\alpha u_\lambda - a \mu \Delta u_\lambda = - \beta \nabla \sigma_\lambda, \\
u_\lambda = 0 \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty, \\
\partial u_\lambda / \partial n = 0 \quad \text{on} \quad \Gamma_\infty^+, \\
\end{cases} \tag{12.15}
\]

\[
\begin{cases}
\alpha^2 \psi_\lambda - \xi \Delta \psi_\lambda = \alpha \sigma_\lambda + \nabla \cdot u_\lambda. \\
\psi_\lambda \quad \text{satisfying} \quad (12.12). \\
\end{cases} \tag{12.16}
\]

Then we define \( A \) by:

\[ A \lambda = \left\{ -\partial \psi_\lambda / \partial n \mid \Gamma_B \cup \Gamma_\infty, \quad \psi_\lambda \mid \Gamma_\infty \right\}. \tag{12.17} \]

**Construction of the explicit part of the boundary integral equation**

Let us define \( \sigma_0, u_0, \psi_0 \) solutions of (12.18), (12.19), (12.20) respectively

\[
\begin{cases}
\alpha^2 \sigma_0 - \xi \Delta \sigma_0 = \alpha g \quad \nabla \cdot f \quad a \mu \Delta g, \\
\sigma_0 = 0 \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty, \\
\partial \sigma_0 / \partial n = 0 \quad \text{on} \quad \Gamma_\infty^+, \\
\end{cases} \tag{12.18}
\]

\[
\begin{cases}
\alpha u_0 - a \mu \Delta u_0 = f - \beta \nabla \sigma_0, \\
u_0 \quad \text{satisfying} \quad (12.5), \\
\end{cases} \tag{12.19}
\]

\[
\begin{cases}
\alpha^2 \psi_0 - \xi \Delta \psi_0 = a \sigma_0 + \nabla \cdot u_0 - g, \quad \psi_0 \quad \text{satisfying} \quad (12.12), \\
\end{cases} \tag{12.20}
\]

and we obtain

\[ b_0 = \left\{ \partial \psi_0 / \partial n \mid \Gamma_B \cup \Gamma_\infty, \quad \psi_0 \mid \Gamma_\infty \right\}. \tag{12.21} \]
Then the value of $\lambda$ is given by the linear integral equation:

$$A\lambda = b_0, \quad (12.22)$$

with

$$\lambda = 0 \quad \text{on} \quad \Gamma_\infty,$$

in order to satisfy the boundary conditions (12.4).

It follows from (12.14)–(12.22) that the solution of the generalized Stokes problem is reduced to $2N + 3$ Poisson problems ($N + 2$ to obtain $\psi_0$, $N + 1$ to obtain $\{\sigma, u\}$ once $\lambda$ is known via the boundary integral solution of (12.22)).

13. Least squares–conjugate gradient solution of the nonlinear subproblems (11.9)

13.1. Variational formulations

At each full step of the operator splitting method (11.8)–(11.10), we have to solve a nonlinear system of the following type

$$\begin{cases} \alpha \sigma + u \cdot \nabla \sigma = g, \\ \alpha u - b \mu \Delta u - \psi(\sigma, u, T) = f, \\ \alpha T - b \Pi \Delta T - \chi(\sigma, u, T) = h, \end{cases} \quad (13.1a)$$

$$\begin{cases} \sigma = \sigma_B \quad \text{on} \quad \Gamma_B, \\ \sigma = 0 \quad \text{on} \quad \Gamma_\infty^-, \end{cases} \quad (13.1b)$$

$$\begin{cases} u = 0 \quad \text{on} \quad \Gamma_B, \\ u = u_\infty \quad \text{on} \quad \Gamma_\infty^-, \end{cases} \quad (13.1c)$$

$$\begin{cases} \partial u / \partial n = 0 \quad \text{on} \quad \Gamma_\infty^+, \end{cases} \quad (13.1d)$$

where $\alpha$ is a positive parameter and where $g$, $f$, $h$ are given functions. We denote by $\sigma_B$ the rigid boundary value of $\sigma$ calculated in the previous Stokes step.

Using the notations of fig. 10.1, we introduce the following functional spaces:

$$R_\tau = \{ \phi \in H^1(\Omega) \mid \phi = r \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty^- \}, \quad (13.2)$$

$$W_z = \{ v \in (H^1(\Omega))^n \mid v = z \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty^- \}, \quad (13.3)$$

$$V_s = \{ \theta \mid \theta \in H^1(\Omega), \theta = s \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty^- \}, \quad (13.4)$$
where $H^1(\Omega)$ is defined by (3.2). If $r$ (resp. $z$, $s$) is sufficiently smooth, then $R_r$ (resp. $W_z$, $V_s$) is non-empty (the above choice for the space of the densities $R_r$ is motivated by the fact that $\rho$ will be approximated by continuous functions and that the restriction of $\rho$ to $\Gamma_{\infty}^-$ makes sense; of course, this supposes implicitly that $\rho$ has some regularity).

Then an equivalent variational formulation of cqs. (13.1) is

$$
\begin{aligned}
\alpha \int_{\Omega} \sigma \phi \, dx + \int_{\Omega} u \cdot \nabla \sigma \phi \, dx = \int_{\Omega} g \phi \, dx, \\
\alpha \int_{\Omega} u \cdot v \, dx + b \mu \int_{\Omega} \nabla u \cdot \nabla v \, dx - \int_{\Omega} \psi(\sigma, u, T) \cdot v \, dx = \int_{\Omega} f \cdot v \, dx, \\
\alpha \int_{\Omega} T \theta \, dx + b \Pi \int_{\Omega} \nabla T \cdot \nabla \theta \, dx - \int_{\Omega} \chi(\sigma, u, T) \theta \, dx = \int_{\Omega} h \theta \, dx.
\end{aligned}
$$

\[
\forall \{\phi, v, \theta\} \in R_0 \times W_z \times V_s, \quad \{\sigma, u, T\} \in R_r \times W_z \times V_s,
\]

where the value of $r$, $z$, $s$ are precised by the boundary conditions of (13.1).

13.2. Least squares formulation

A natural least squares formulation of (13.5) is

$$
\begin{aligned}
\text{Find } \{\sigma, u, T\} & \in R_r \times W_z \times V_s \text{ such that } \\
J(\sigma, u, T) & \leq J(\eta, w, \tau), \forall \{\eta, w, \tau\} \in R_r \times W_z \times V_s,
\end{aligned}
$$

with

$$
J(\eta, w, \tau) = \frac{\alpha}{2} \int_{\Omega} |\sigma - \eta|^2 \, dx + \frac{A}{2} \left( \alpha \int_{\Omega} |u - w|^2 \, dx + b \mu \int_{\Omega} |\nabla (u - w)|^2 \, dx \right) \\
+ \frac{B}{2} \left( \alpha \int_{\Omega} |T - \tau|^2 \, dx + b \Pi \int_{\Omega} |\nabla (T - \tau)|^2 \, dx \right).
$$

where, in (13.7), we have $A > 0$, $B > 0$, and where $\sigma, u, T$ are functions of $\{\eta, w, \tau\}$ through the state equations:

$$
\begin{aligned}
\alpha \int_{\Omega} \sigma \phi \, dx = \int_{\Omega} g \phi \, dx - \int_{\Omega} w \cdot \nabla \eta \phi \, dx, & \quad \forall \phi \in R_0, \\
\alpha \int_{\Omega} u \cdot v \, dx + b \mu \int_{\Omega} \nabla u \cdot \nabla v \, dx = \int_{\Omega} \psi(\eta, w, \tau) \cdot v \, dx + \int_{\Omega} f \cdot v \, dx, & \quad \forall v \in W_z, \\
\alpha \int_{\Omega} T \theta \, dx + b \Pi \int_{\Omega} \nabla T \cdot \nabla \theta \, dx = \int_{\Omega} \chi(\eta, w, \tau) \theta \, dx + \int_{\Omega} h \theta \, dx, & \quad \forall \theta \in V_s.
\end{aligned}
$$

In order to solve the above minimization problem (13.6)-(13.8), we use a conjugate gradient algorithm analogous to the one described in section 3.4 for the incompressible case.
13.3. Conjugate gradient solution of the least squares problem (13.6)

The space $\mathbb{R}_x \times \mathbb{W}_x \times \mathbb{V}_x$ is equipped with the product norm below:

$$
\{ \phi, v, \theta \} \rightarrow \left\{ \alpha \int_\Omega \phi^2 \, dx + \alpha A \int_\Omega v^2 \, dx + b \mu A \int_\Omega |\nabla v|^2 \, dx \\
+ \alpha B \int_\Omega \theta^2 \, dx + b II B \int_\Omega |\nabla \theta|^2 \, dx \right\}^{1/2},
$$

then, the algorithm can be written as follows:

**Step 0: Initialization**

$$
\{ \eta^0, w^0, \tau^0 \} \in \mathbb{R}_x \times \mathbb{W}_x \times \mathbb{V}_x \text{ given};
$$

(13.9)

solve then the following problems:

$$
\alpha \int_\Omega g^0_\eta \phi \, dx = \left\langle \frac{\partial J}{\partial \eta} (\eta^0, w^0, \tau^0), \phi \right\rangle \quad \forall \phi \in \mathbb{R}_x; \quad g^0_\eta \in \mathbb{R}_x,
$$

(13.10)

$$
\alpha \int_\Omega g^0_w \cdot v \, dx + b \mu \int_\Omega \nabla g^0_w \cdot \nabla v \, dx = \left\langle \frac{\partial J}{\partial w} (\eta^0, w^0, \tau^0), v \right\rangle \quad \forall v \in \mathbb{W}_x; \quad g^0_w \in \mathbb{W}_x,
$$

(13.11)

$$
\alpha \int_\Omega g^0_\tau \theta \, dx + b II \int_\Omega \nabla g^0_\tau \cdot \nabla \theta \, dx = \left\langle \frac{\partial J}{\partial \tau} (\eta^0, w^0, \tau^0), \theta \right\rangle \quad \forall \theta \in \mathbb{V}_x; \quad g^0_\tau \in \mathbb{V}_x,
$$

(13.12)

where $\langle \cdot, \cdot \rangle$ denotes duality pairings, and set

$$
z^0_\eta = g^0_\eta, \quad z^0_w = g^0_w, \quad z^0_\tau = g^0_\tau.
$$

(13.13)

Then for $n \geq 0$, assuming $\{ \eta^n, w^n, \tau^n \}, \quad g^n = \{ g^n_\eta, g^n_w, g^n_\tau \}, \quad z^n = \{ z^n_\eta, z^n_w, z^n_\tau \}$ known, we obtain $\{ \eta^{n+1}, w^{n+1}, \tau^{n+1} \}, \quad g^{n+1} = \{ g^{n+1}_\eta, g^{n+1}_w, g^{n+1}_\tau \}, \quad z^{n+1} = \{ z^{n+1}_\eta, z^{n+1}_w, z^{n+1}_\tau \}$ as follows:

**Step 1: Descent**

$$
\text{Find } \lambda_n \in \mathbb{R}_x \text{ such that } \\
J(\{ \eta^n, w^n, \tau^n \} - \lambda_n z^n) \leq J(\{ \eta^n, w^n, \tau^n \} - \lambda z^n) \quad \forall \lambda \in \mathbb{R}_x,
$$

(13.14)

and set

$$
\{ \eta^{n+1}, w^{n+1}, \tau^{n+1} \} = \{ \eta^n, w^n, \tau^n \} - \lambda_n z^n.
$$

(13.15)
Step 2: Calculation of the new descent direction

Solve

\[ \alpha \int_{\Omega} g_{\eta}^{n+1} \phi \, dx = \left( \frac{\partial J}{\partial \eta} (\eta^{n+1}, w^{n+1}, \tau^{n+1}), \phi \right) \forall \phi \in \mathbb{R}_n; \, g_{\eta}^{n+1} \in \mathbb{R}_n. \]  

(13.16)

\[ \alpha \int_{\Omega} g_{w}^{n+1} \cdot \mathbf{v} \, dx + b \mu \int_{\Omega} \nabla g_{w}^{n+1} \cdot \nabla \mathbf{v} \, dx = \left( \frac{\partial J}{\partial w} (\eta^{n+1}, w^{n+1}, \tau^{n+1}), \mathbf{v} \right) \forall \mathbf{v} \in \mathbb{W}_0; \, g_{w}^{n+1} \in \mathbb{W}_0. \]

(13.17)

\[ \alpha \int_{\Omega} g_{\tau}^{n+1} \theta \, dx + b \Pi \int_{\Omega} \nabla g_{\tau}^{n+1} \cdot \nabla \theta \, dx = \left( \frac{\partial J}{\partial \tau} (\eta^{n+1}, w^{n+1}, \tau^{n+1}), \theta \right) \forall \theta \in \mathcal{V}_0, \, g_{\tau}^{0} \in \mathcal{V}_0. \]

(13.18)

Compute then

\[ \gamma_n = \left( \left( g_{\eta}^{n+1}, g_{w}^{n+1}, g_{\tau}^{n+1} \right), \left( g_{\eta}^{n+1} - g_{\eta}^{n}, g_{w}^{n+1} - g_{w}^{n}, g_{\tau}^{n+1} - g_{\tau}^{n} \right) \right), \]

(13.19)

where \( \forall \{ \phi, \mathbf{v}, \theta \}, \{ \phi', \mathbf{v}', \theta' \} \), we have

\[ \left( \{ \phi, \mathbf{v}, \theta \}, \{ \phi', \mathbf{v}', \theta' \} \right)_1 = \alpha \int_{\Omega} \phi \phi' \, dx + \alpha A \int_{\Omega} \mathbf{v} \cdot \mathbf{v}' \, dx + b \mu A \int_{\Omega} \nabla \mathbf{v} \cdot \nabla \mathbf{v}' \, dx \]

\[ + \alpha B \int_{\Omega} \theta \theta' \, dx + b \Pi B \int_{\Omega} \nabla \theta \cdot \nabla \theta' \, dx. \]

Finally define \( z^{n+1} \) by

\[ z^{n+1} = g^{n+1} + \gamma_n z^n. \]

(13.20)

Do \( n = n + 1 \) and go to (13.14).

Applying algorithm (13.9)-(13.20) to solve the least squares problem (13.6) requires the solution at each iteration of several decoupled Poisson problems associated to an elliptic operator of the type \( (\alpha I - \mu \Delta) \).

We have also tested some of these methods combining the features of conjugate gradient with updated variable preconditioning and quasi-Newton algorithm (see refs. [31,32]) which have proved performing (20\% faster) and robust for the solution of (13.6).

14. Finite element approximation

For simplicity, we suppose that \( \Omega \) is a bounded polygonal domain of \( \mathbb{R}^2 \) but the method is easily extended to domain of \( \mathbb{R}^3 \). The space discretization is achieved by a finite element
approximation. With $T_h$ a standard finite element triangulation of $\Omega$, and $h$ the maximal length of the edges of the triangles of $T_h$, we introduce the following discrete spaces (with $P_k$ = space of polynomials in two variables of degree $\leq k$):

\begin{align}
R_h &= \left\{ \phi_h : \phi_h \in C^0(\bar{\Omega}), \quad \phi_h|_T \in P_1, \quad \forall T \in T_h, \quad \phi_h - r_h \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty \right\}, \quad (14.1) \\
W_h &= \left\{ \omega_h : \omega_h \in \left( C^0(\bar{\Omega}) \right)^2, \quad \omega_h|_T \in P_1 \times P_1, \quad \forall T \in T_h, \quad \omega_h = \alpha_h \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty \right\}, \quad (14.2) \\
V_h &= \left\{ \theta_h : \theta_h \in C^0(\bar{\Omega}), \quad \theta_h|_T \in P_1, \quad \forall T \in T_h, \quad \theta_h = \beta_h \quad \text{on} \quad \Gamma_B \cup \Gamma_\infty \right\}. \quad (14.3)
\end{align}

Then, we can write for the discrete problem, the algorithm previously defined for the continuous one; concerning the construction of the approximate integral operator $A_h$, we refer to Glowinski [1], Péraux [73].

We observe that unlike in the approximation of the incompressible Navier–Stokes equations, we use discrete spaces of the same type to approximate $\sigma$, $u$ and $T$.

15. Self-adaptive mesh refinement techniques for accurate solution of the compressible Navier–Stokes equations

The study presented in this section is based on the recent works of Palmerio [74] and Pouletty [75].

The geometric flexibility of triangular elements is well suited for refinement processes. The basic idea is, starting from a solution on an initial mesh, to improve locally the accuracy of this solution by refining the mesh according to criteria which are functions of the known solution. Then the mesh is depending not only on the geometry or on an a priori knowledge of the critical areas of the computational domain, but is also influenced a posteriori by the solution itself.

Presently, we are considering adaptive mesh refinement only for steady state problems, the generalization to transient problems is a most interesting question, but it still requires further investigations.

The refinement techniques are based on two procedures:

i) moving nodes, their number remaining constant [76,77],

ii) adding nodes, the location of the old ones being fixed [74,78].

Concerning the criterion, it can be of two types, either numerical based on a posteriori error estimates [79,80] or physical [78,81] (areas of large gradients, for instance).

The technique presented here is based principally on (ii) and physical criteria.

15.1. Global algorithm

We proceed via the following algorithm:

(i) Let $T_h^0$ be the initial triangulation (or tetrahedrization in 3D) and $u^0$ a quasi-converged solution of the considered problem.

Then, for $k \geq 0$, knowing the solution $u^k$ on the triangulation $T_h^k$, we define $T_h^{k+1}$, $u^{k+1}$ by:
(ii) Construction of $E^k$, set of the elements to be divided.

Let $C^k(T)$ be the value of the criterion, $T \in \mathbb{T}_h^k$. If

$$C^k(T) > \varepsilon \max_{T \in \mathbb{T}_h^k} C^k(T)$$

then

$$T \in E^k$$

with

$$0 < \varepsilon < 1,$$

to be specified.

(iii) Construction of the new triangulation $\mathbb{T}_h^{k+1}$.

(iv) Computation of $u^{k+1}$ on $\mathbb{T}_h^{k+1}$.

The initialization of $u^{k+1}$ is given by linear interpolation of $u^k$ on $\mathbb{T}_h^{k+1}$.

Let us now describe in details the important steps (ii) in section 15.3 and (iii) in section 15.2, of the above algorithm.

15.2. Local refinement technique, construction of $T_h^{k+1}$ knowing $E^k$

For the areas containing the elements of $E^k$, the aim is to replace the elements of size $h$ by elements of size $h/2$, ensuring moreover a good quality of the triangulation.

We give thereafter details for the 2D case (the 3D refinement is based on the same methodology, but introduces some extra geometrical difficulties (cf. remark (15.2))):

(i) Each triangle $T$ of $E^k$ is divided into four sub-triangles by adding a node at each mid-side; then a triangle $T$ of the previous triangulation $\mathbb{T}_h^k$ has $j$ nodes added ($j = 0, \ldots, 3$) (cf. fig. 15.1).

(ii) If $j = 1, \ldots, 3$, the triangles have to be sub-divided in order to obtain a conforming finite element triangulation, and, in the matching area ($j = 1, 2$), special care is required in order to ensure a “good quality” of the resulting triangulation. Therefore, we apply the following rules:

- $j = 0$, triangle $T$ unchanged,
- $j = 1$, one node added, triangle $T$ divided into two sub-triangles,
- $j = 2$, if $T$ is divided into 3 sub-triangles, their shape is considered as unacceptable, then a third node is added and $T$ is divided into four elements.
- $j = 3$, $T$ is divided into 4 sub-triangles.

Fig. 15.1.
Iterative use of the above procedure generates a new mesh.

In order to maintain a good level of quality of the final triangulation, the following remarks are taken into account:

**Remark 15.1.** If a triangle with 2 vertices on the body has one node added, then two extra nodes are added and the triangle is divided into four sub-triangles.

**Remark 15.2.** If a node is added on a side adjacent to the body, this point is projected on the body (either by a parabolic interpolation or, if possible, by using the analytical function defining the body). For 3D industrial application, this technique requires rather important geometrical tools (interpolation, splines) in order to guarantee that new nodes are smoothly projected on the body.

**Remark 15.3.** In the areas where triangles have been sub-divided, a local elasticity-like equation [75] is applied in order to move locally the nodes and to obtain a smoother mesh. This technique is particularly efficient in the matching region where triangles have been sub-divided into two sub-triangles.

Among other moving nodes techniques to improve the smoothing of the mesh, we want to mention the recent works of Cua [82] using generalized gradient methods.

**Remark 15.4.** The quality of the triangulation is also related to the number of neighbors surrounding a node which has to be upper limited. This restriction applies to the triangles divided into two and to the node from which the splitting lines are issued. If more than two of them are issued from the same node, this situation is automatically rejected and in this case, the peculiar triangles are subdivided into four (see fig. 15.2).

In practice, the matching region connecting refined elements of $E_k$ to non-refined elements of $T_h^k$ is restricted to a limited size strip.

### 15.3. Choice of criteria for the compressible Navier–Stokes equations

For a viscous flow simulation, it is important to refine the mesh in the vicinity of the shocks and in the boundary layer and wake regions. For this reason, we introduce different criteria which can be combined:

\[
C_1(T) = |u \cdot \nabla M|_T, \tag{15.1}
\]

\[
C_2(T) = |u \cdot \nabla M|_{T/\|u\|_T}, \tag{15.2}
\]
Fig. 15.3. Adapted meshes obtained with different criteria for the simulation of the flow around a NACA 0012, $M_\infty = 0.8$; Re = 500; $\alpha = 10^\circ$. (a1) $C_1(T)$; 2689 nodes, 5221 elements; (a2) $C_2(T)$; 2624 nodes, 5085 elements; (b) $C_3(T)$; 2652 nodes, 5122 elements; (c) $C_4(T)$; 2687 nodes, 5186 elements; (d1) $C_6(T)$; 2664 nodes, 5150 elements; (d2) $C_7(T)$; 2657 nodes, 5131 elements; (d3) $C_1 + C_6$; 2630 nodes, 5083 elements; (d4) $C_2 + C_7$; 2623 nodes, 5069 elements.
\[ C_3(T) = \| \nabla M \|_T, \]  
\[ C_4(T) = \| \nabla \times u \|_T, \]  
\[ C_5(T) = \| \nabla \times u \|_T / \| u \|_T, \]  
\[ C_6(T) = \| u \times \nabla M \|_T, \]  
\[ C_7(T) = \| u \times \nabla M \|_T / \| u \|_T, \]  
\[ C_8(T) = C_7(T) \times L_T, \quad \forall T \in T_h, \]  

using averaging formulae if necessary, these different criteria are defined as piecewise constant on each triangle (or tetrahedron) \( T \) of \( T_h \).

In (15.1)–(15.8):

(a) \( M \) is the local Mach number, \( M = \| u \| / c \), with \( c = \sqrt{\gamma p / \rho} \).

(b) \( L_T \) is the distance between the triangle under consideration and the leading edge.

Several combinations of these criteria acting in preferred directions have been tested to generate adapted refined grids for the computation of a viscous flow around a NACA0012 airfoil as depicted in fig. 15.3.

Self-adaptive mesh refinement techniques provide for a given number of nodes, more accurate results (cf. ref. [81]). In order to improve efficiency of the algorithm, the step under investigation is to develop a full multigrid method on the different adapted meshes [83,86].

16. Numerical experiments

We illustrate the methods described in the above section by the presentation of the results of numerical experiments, where these methods have been applied to simulate some compressible viscous 2D and 3D flows of practical interest. All the calculations which follow have been performed either on IBM 3090 of AMD-BA industries or on CRAY 1S of CCVR using the finite element method associated to the discrete spaces defined by (14.1)–(14.3).

16.1. A first class of test problems

We consider the solution of the Navier–Stokes equations for the flow of compressible viscous fluids around a NACA0012 airfoil (fig. 16.1). We have selected this problem since it is a quite classical and significant test problem for Navier–Stokes solvers [87]. The method described in the above sections 9–14 has been applied to compute the steady solution of (10.5)–(10.7) for the two following tests.

The first test is a 2D transonic calculation at \( M_\infty = 0.85, \) \( Re = 500, \) at an angle of attack of \( \alpha = 0^\circ. \) Fig. 16.1 shows an enlargement of the triangulation near the profile and fig. 16.2 shows the computed pressure coefficient. Results of fig. 16.3, Mach distribution and density distribution have been obtained with scheme (11.8)–(11.10) after 50 time cycles (time step \( \Delta t = 0.2 \)).
The second test is a 2D supersonic flow around a NACA0012 at $M_\infty = 2$, $Re = 10^6$, at an angle of attack of $10^\circ$. Fig. 16.4 shows an enlargement of the locally refined triangulation generated by self-adaptive algorithms (see section 16.3.1). In fig. 16.5 is represented the pressure distribution on the airfoil obtained with scheme (11.8)-(11.10). The colored pictures of fig. 16.6 (see page 159) show the pressure, Mach and density distributions around the airfoil. These results are in good agreement with those obtained by other Navier–Stokes solvers [87].
Fig. 16.3. Flow around a NACA 0012 airfoil, $M_\infty = 0.85, \text{Re} = 500$. (a) Mach contours; (b) density contours.
Fig. 16.5. Pressure coefficient on the NACA 0012 airfoil, $M_a = 2; \ Re = 106; \ \alpha = 10^\circ$. 

Fig. 16.4. Adapted mesh, $M_a = 2; \ Re = 106; \ \alpha = 10^\circ$. 

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Fig. 16.6. Flow around a NACA 0012 airfoil. $M_\infty = 2; \text{Re} = 106; \alpha = 10^\circ$. (a) Pressure map; (b) Mach number map; (c) density map.

Fig. 16.21. Vorticity maps. $M_\infty = 0.9, \text{Re} = 100, \alpha = 40^\circ$. (a) Vertical section; (b) cross section; (c) on the inlet.
Fig. 16.7. Comparison of density contours. NACA 0012 airfoil, $M_\infty = 2$; $Re = 106$; $\alpha = 10$. (a) Computational results; (b) experimental results.
Fig. 16.8. Velocity distribution, $M_\infty = 0.7$; $Re = 100$; $\alpha = 30^\circ$; $\Delta t = 0.2$. (a) $t = 8$; (b) $t = 12$; (c) $t = 16$; (d) $t = 20$; (e) $t = 24$; (f) $t = 28$; (g) $t = 32$; (h) $t = 36$. 
Fig. 16.8 (continued).
Fig. 16.8 (continued).
Fig. 16.8 (continued).
For this case, a comparison between numerical results and measurements due to Allegre et al. [88] has been done. In fig. 16.7, we have plotted and compared the density lines obtained with the numerical method and the experiments.

It can be assumed that the computational results would be still in better agreement with measurements if the adherence condition on velocity prescribed at the rigid boundaries, was replaced by some velocity distribution taking more closely into account the boundary layer phenomenon.

Finally, we have also considered an unsteady problem around the same NACA0012 airfoil with the following data: $M_{\infty} = 0.7$, $Re = 100$, $\alpha = 30^\circ$. A time dependent solution has been obtained starting from a free-stream initial solution. Fig. 16.8 describes the vortex pattern of the flow at different time steps ($t = 8, 12, 16, \ldots, 36$, selected time step $\Delta t = 0.2$) on which we can observe the creation and propagation of vortices in the deviated wake of the airfoil.

16.2. A second class of problems

A second class of problems that we consider is much more complicated than the first one, since it concerns the unsteady simulation of a transonic viscous flow around and inside a 2D nozzle at high incidence $\alpha = 30^\circ$, mach number $M_{\infty} = 0.9$ and Reynolds number $Re = 100$ (the characteristic length being the distance between the walls of the nozzle). Fig. 16.9 shows the details of an adapted grid close to the air intake obtained with adaptive mesh refinement techniques (section 15.). Figs. 16.10, 16.11 show the Mach lines and the vortex pattern at different time cycles ($t = 16, 20, \ldots, 44$). We can observe the creation of vorticity at the

Fig. 16.9. Adapted mesh around and inside an inlet, 3811 nodes, 7303 triangles in the global mesh.
Fig. 16.10. Flow around and inside an inlet. Mach lines, $M_{\infty} = 0.9$, $Re = 100$; $\alpha = 30^\circ$; $\Delta t = 0.2$. (a) $t = 16$; (b) $t = 20$; (c) $t = 24$; (d) $t = 28$; (e) $t = 32$; (f) $t = 36$; (g) $t = 40$; (h) $t = 44$. 
Fig. 16.10 (continued).
Fig. 16.10 (continued).
Fig. 16.10 (continued).
Fig. 16.11. Flow around and inside an inlet, velocity distribution. $M_\infty = 0.9$; $Re = 100$; $\alpha = 30^\circ$; $\Delta t = 0.2$ (a) $t = 16$; (b) $t = 20$; (c) $t = 24$; (d) $t = 28$; (e) $t = 32$. (f) $t = 36$; (g) $t = 40$; (h) $t = 44$. 
Fig. 16.11 (continued).
Fig. 16.11 (continued).
leading edge and the growing structures which propagate around the inlet. A global view of the deviated wake behind the inlet is shown in fig. 16.12.

16.3. 2D and 3D Navier–Stokes solutions using self-adaptive mesh refinements

The methodology described in section 15 has been applied to various numerical 2D and 3D simulations around geometries of industrial interest in order to improve unexpensively local accuracy of the Navier–Stokes solutions.

16.3.1. Local enrichment around a NACA0012

The main features of a viscous flow are difficult to identify due to its complexity: physical criteria in preferred directions have to be found to localize not only discontinuities but also boundary layers and wakes and a very important work remains still open in the broad area of unsteady flow with moving features. In the following, we will concentrate only on steady solutions. We consider the numerical simulation of a supersonic flow at $M_\infty = 2$, $Re = 10^6$, at an angle of attack of $\alpha = 10^\circ$ (section 16.1). Starting from a computation on a coarse mesh (800 nodes, 1 514 triangles), two successive refined meshes are generated as shown in fig. 16.13 using criterion $C_1 + C_6$ (see section 15.3). Comparison of the corresponding Mach contours (fig. 16.14) shows a sensitive improvement of the accuracy with an increasing stability of the lines and also a better capture of the shock layer.

16.3.2. 3D Navier–Stokes solution around an inlet

Finally a 3D computation of industrial interest around and inside an idealized air intake at $M_\infty = 0.9$, $Re = 100$, and $\alpha = 40^\circ$, using local enrichment by self adaptive procedure has been also performed. The initial mesh is depicted in figs. 16.15a–16.17a by a view of the inlet and by vertical sections in the flow direction and perpendicular to the flow direction. The mesh is locally
enriched using criterion $C_4$, an exceeded threshold of the vorticity generates finer elements in the vicinity of the leading and trailing edges of the air intake. After two successive enrichments, the adapted mesh depicted by figs. 16.15b–16.17b is obtained. The pressure and Mach contours of the solution computed on the initial coarse mesh are shown in fig. 16.18a, 16.19a, while the same iso lines of the solution on the refined mesh are presented in figs. 16.18b, 16.19b. A significant improvement of the quality of pressure and Mach lines can be observed specially in the vicinity
Fig. 16.14. Mach contours on adapted meshes, $M_\infty = 2$; $Re = 10^6$, $\alpha = 10^\circ$. (a) Nodes: 1797, elements: 3474; (b) nodes: 2646, elements: 5172.
Fig. 16.15. Details of the initial and adapted meshes on a 3D inlet.
Fig. 16.16. Initial and adapted meshes around and inside a 3D inlet. Vertical section in flow direction. (a) Initial mesh, nodes: 3096, elements: 16632; (b) adapted mesh: nodes: 6958, elements: 34760.
Fig. 16.17. Vertical section perpendicular to the flow direction.
Fig. 16.18. Pressure lines around and inside a 3D inlet. Vertical section in the flow direction. \( M_\infty = 0.9, \) \( Re = 100, \) \( \alpha = 40^\circ. \)
Fig. 16.19. Mach lines around and inside a 3D inlet. Vertical section in the flow direction. $M_{\infty} = 0.9$, $Re = 10^6$, $\alpha = 40^\circ$. 
Fig. 16.20. Velocity distribution. Vertical section in the flow direction. $M_\infty = 0.9$, $Re = 100$, $\alpha = 40^\circ$. 
of the inlet entrance and also in the boundary layer regions. Fig. 16.20 shows the velocity distribution in a vertical section, obtained on the finest grid and in fig. 16.21 (see page 159) are presented different views of the vorticity distribution. As a solution has been now obtained with the above enrichment procedure, the next step will be to accelerate its convergence via adapted full multigrid techniques.

Conclusion

We have presented in this paper a family of methods for solving viscous flow problems modeled by the Navier–Stokes equations. Efficient and accurate methods exist now for incompressible flow simulations on practically any geometry, at Reynolds numbers of the order of $10^3$ to $10^4$, in two and three dimensions. Recent progresses suggest that a similar situation may hold very soon concerning compressible flow problems.

However a lot of theoretical and computational work has still to be done in order to extend the above methods to turbulent and/or reacting flows.

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References

Equations, Lecture Notes at the Von Karman Institute, March 1986, Report HE41/86.03, Electricité de France, Chatou.


